#### APPENDIX A

# Pre-processing of the Free-Air Gravity Anomalies (FAGA)

## A.1 Data Description

possibility for examining heretofore unknown aspects of a remote region of the Earth. been widely distributed and used in many previous investigations, they offer a real offshore with an average trackline spacing of 30 km. Because these data have not the central ice-covered region, above the coastal areas and out approximately 100 km ical Projects of 1991 and 1992 (GAP91 and GAP92). These data were collected over by Brozena [1995] for the Naval Research Laboratory for the Greenland Aerogeophysto generate this grid originate from several surveys, including observations processed (NIMA) for the Greenland area  $(58.7 - 84.2^{\circ}N)$  and  $285.75 - 349.50^{\circ}E$ ). The data used A FAGA grid has been determined by the National Imagery and Mapping Agency

tion errors observations to generate 5 arcminute grids of FAGA values (Figure A.1) and observa-(DEM) to facilitate comparisons with surface control measurements (Forsberg and Kenyon, 1995). Predictions were made on this digital elevation model NIMA used least squares collocation on terrestrial, marine and airborne gravity (Figure A.2) at the respective elevations (Figure A.3) for each grid point

FAGA field contains broader longer wavelength features over the smoother, central, Examination of Greenland's coastal regions in Figure A.1 shows the severe effects being located upon the varying surface defined in Figure

regions. levels and subglacial crustal sources for the interior region as compared to the coastal the ice sheet. This in part may be due to the further distance between the observation ice-covered regions and shorter wavelength features nearer the coasts and the edge of

to the offshore regions of Northeast and Southeast of Greenland (Figure A.2). Labrador Straits. Most of the areas containing larger observation errors are confined for those portions of the grid that are onshore or over enclosed waterways such as the The errors associated with these measurements indicate a high degree of confidence

by removing terrain effects and emphasizing the remaining FAGA components elevation and not on the DEM's surface. This will facilitate potential field analysis the purposes of this analysis, it is necessary to have all values at a uniform

that would inhibit the total calculation. close to the regions where gravity predictions are desired, singularities may develop several nodes to emulate the gravity effect of the entire modeled to determine their gravity effects. For each prism, densities are assigned to accuracy of the inversion. In GLQ, prisms of material determined from a DEM are nodes used by is also desirable to have the FAGA values sufficiently high enough above the Gaussian Legendre Quadrature integration (GLQ) to optimize prism. If the nodes are to

data, altimeter-derived FAGA, and terrain DEM's). data set to which these gravity data will be directly or indirectly compared (magnetic average data spacing for observed gravity data sets, as well as the intervals of other (longitude) grid or about 10 km by 10 km. This gridding interval was based upon the Finally, the desired data spacing for grids was selected at 6' (latitude) by 15'

## A.2 Upward Continuation

above the source region inside the Earth. of the remaining signal lowered to emulate the effects of data being observed further This would result in wavelengths below 20 km being de-emphasized and the power data were now upward continued to 20 km using the same software (Figure A.4). the authors of the original FAGA data [Forsberg and Kenyon, 1995]. Therefore, these elevations to the DEM surface (Figure A.3) using GEOFOUR software developed by The original FAGA grid (Figure A.1) had been downward continued from aircraft An ideal elevation would be double the selected data spacing or about 20 km.

when examining smaller offshore features. in Figure A.1 were originally at 0 km elevation, therefore, they may be of use later and longer wavelength information (CC=0.87). Gravity data for the offshore regions Figure A.1 and Figure A.4 shows that they essentially contain the same intermediate information contained in the upward-continued data has occurred. of the original data (39 mgals). Therefore, I assume that no serious degradation in the The energy of the upward-continued data (26 mgals) can be seen to be two thirds Examination of

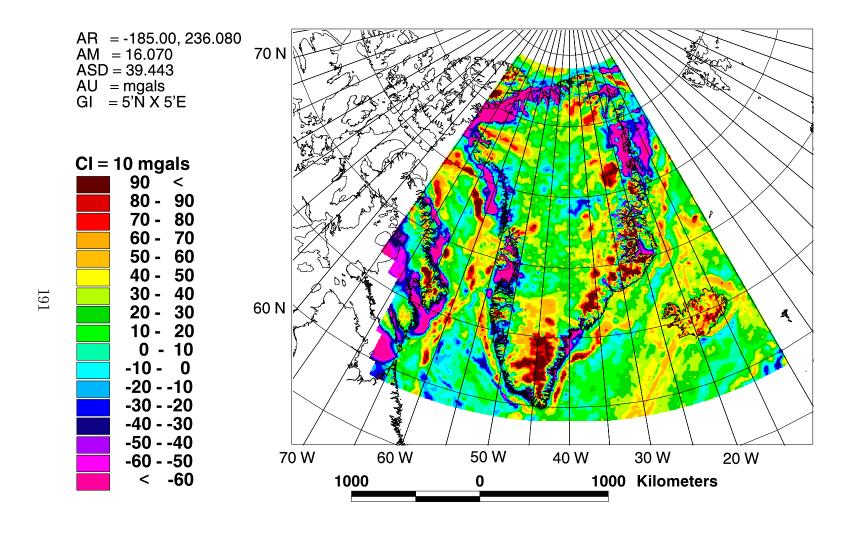


Figure A.1: Original NIMA Free-Air Gravity Anomalies (FAGA) at the Earth's surface shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W. Data are registered to variable surface elevations that are given in Figure A.3.

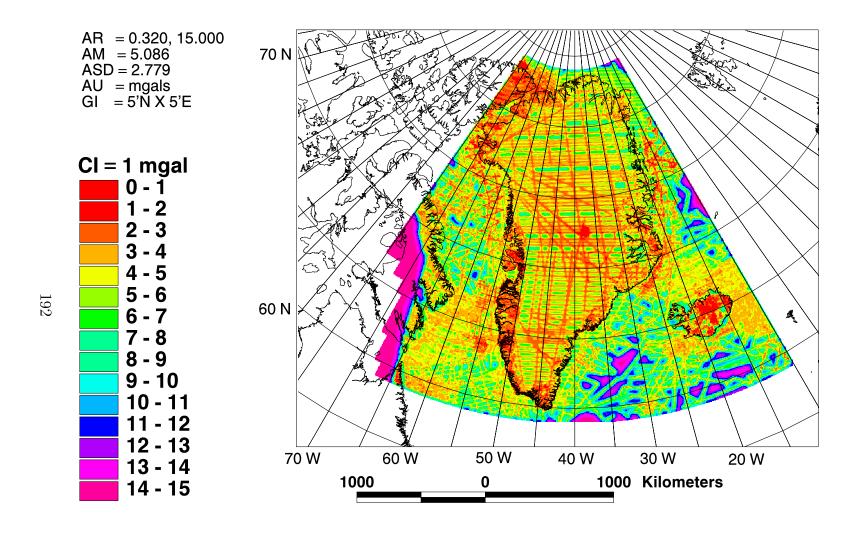


Figure A.2: Observation errors provided with the NIMA FAGA shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W.

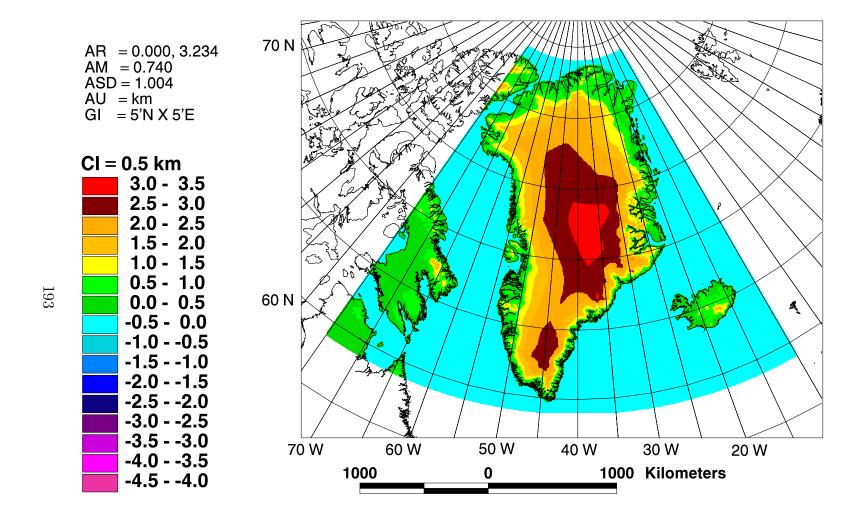


Figure A.3: Orthometric heights associated with the NIMA FAGA above mean sea level shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W.

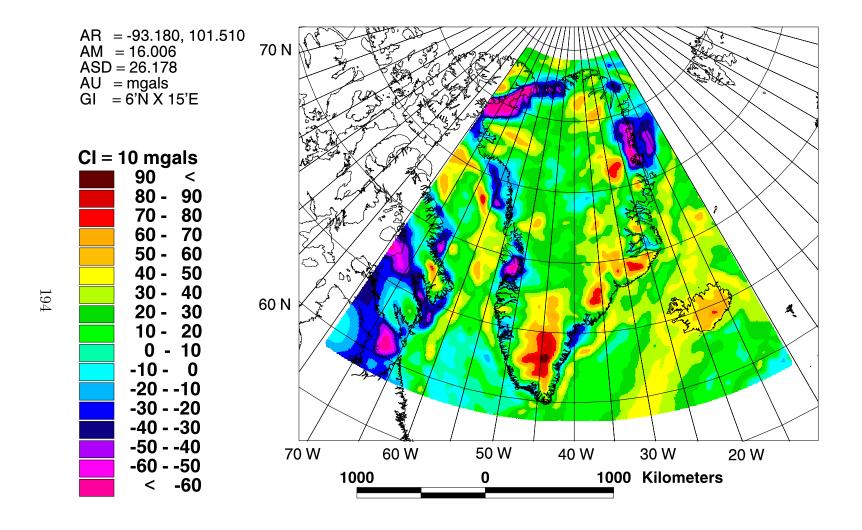


Figure A.4: NIMA FAGA upward continued to Z=20 km elevation shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W.

#### APPENDIX B

## Pre-processing of the Magnetic Anomalies (MA)

## **B.1** Data Description

and North America. veys covering the eastern Arctic Ocean and most other land areas, including Eurasia and GAP92). Additional airborne surveys covered the western Arctic Ocean and the data set was derived from numerous airborne surveys covering Iceland and Greenfrom several different surveys on both the North Atlantic and the East and West sources of data: shipborne, airborne, and prior gridded data sets. entire Arctic and North Atlantic region in a 5 km grid, including the Greenland area mation (Figure B.1) in open file 3125b [Verhoef et al., 1996]. These data cover the Grand Banks. (NRL) during the Greenland Aerogeophysical Projects of 1991 and 1992 (GAP91 land, including data collected by Brozena [1995] for the Naval Research Laboratory coasts of Greenland were merged to produce one major data set. Another major (58.7 -84.2° N and 285.75 - 349.50° E). These data were derived from three general The Geological Survey of Canada (GSC) compiled magnetic anomaly (MA) infor-The final source of data derived from several previously gridded sur-Shipborne data

and gridded. To facilitate merging these sources, high pass filtering at 400 km was Extensive efforts were made to merge these disparate data sources into a cohesive Each of the three subsets (ship, air, and grid) were independently merged

data sets passed wavelengths as long as 1000 km at a much reduced power. The final gridding grid is composed of 400 km and shorter wavelengths, although the taper of the filter selected rather than trying to remove individual tilts and biases. Hence, the resulting was derived using minimum curvature from 7.5 million data points in all three major

Geology compilation. the Diskó Island region and was a 2 km grid taken from a Decade of North American the primary sources. The only previously gridded data set that was included covered oceanic areas along the eastern and western coasts, shipborne measurements provide aeromagnetic data plus data derived primarily from the Danish Geological Survey MA determined for the Greenland field area. These included the GAP91 and GAP92 Of all of these data, only a few represent that subset that directly influenced the southern Greenland and off the East coast [Verhoef et al., 1996]. For deeper

be characterized by wavelengths shorter than 400 km. magnetic characteristics. crust (depths to 10 km) or continental crust (depths to 60 km) [Shive et al., 1992]. that magnetic sources will originate above the Curie isotherm in either the oceanic adequately reflected in the 400 km and shorter wavelength components. It is assumed crustal MA sources will have a shorter wavelength characteristic, which should be Below the Curie isotherm, the rocks become randomly magnetized and lose their The filtering at 400 km wavelength is still sufficient for a crustal analysis because With sources no deeper than 60 km, the resulting MA will

have some expression in the final data wavelengths up to 1000 km, the longer wavelengths associated with these sources will some components that are longer than 400 km. Because the hi-pass filter taper passed For broad, shallow source regions such as the Caledonian Orogeny, there may be

## **B.2** Upward Continuation

do represent separate data. these gridded data sets were not included with the other airborne data sources and were typical of those for the above airborne surveys. The aeromagnetic data used for rived primarily from other aeromagnetic surveys occurring at varying elevations that km, while shipborne observations were all at sea level. The gridded data were decollected at elevations between 200 to 400 m with some values were around 1 and 4 elevations provided and not reduced to a common elevation. Most airborne data were GSC merged all the data into the final grid, the various data sets were merged at the ure B.2) for several reasons. The upward continuation was necessary because when The original MA (Figure B.1) were upward continued to 20 km elevation (Fig-

to place them sufficiently above its source region. for comparison with the MA. The FAGA data had been upward continued to 20 km radial derivative of the FAGA may be used to generate pseudo-magnetic anomalies data via Poisson's Relation. By placing both data sets at the same elevation, the 20 km was selected to facilitate comparison with the 20 km upward continued FAGA emphasize these contrasts as 20 km is much greater than any of the elevations at differences in observation elevations. Upward continuation to 20 km will help to dewhich the various airborne, shipborne or gridded data were determined. The value of Amplitudes may vary from one region to the next because of these significant

## **B.3** Reduction to Pole

horizontal and one vertical. Sources located at the geomagnetic poles will have only the observed MA data. The MA can be broken down into three components: two the magnetic pole. To be useful for comparison to the FAGA, the MA data must be effectively reduced MA sources react to the Earth's magnetic field to

geomagnetic pole, or reduced-to-pole (RTP). a vertical field present. For comparison with FAGA data, which is oriented vertically, MA data for all points must be reduced as if it were being observed at the

20 year span between 1970 and 1990 during which the corefield changed over time. declination [Alsdorf et al., 1994]. However, a bulk of the data were collected over This may be easily accomplished spectrally for uniform magnetic inclination and

1985, and 1990 and intensity values generated from the IGRF models for epochs 1970, 1975, 1980, for the Earth's magnetic field components (declination, inclination, and intensity) at year intervals. The International Geomagnetic Reference Field (IGRF) models provide references Figure B.3 was compiled by averaging the declination, inclination,

Northwest. This is an order of magnitude greater than the 1.1° RMS difference As an example, the declination varies from  $-15^{\circ}$  in the Southeast to  $-90^{\circ}$  in the differences obtained by comparing one portion of the field area against another. changing corefield over time will be assumed negligible compared to the magnitude of relatively small compared to the magnitude of the average declination and inclination of 1.1° and the inclination was characterized by a RMS difference of 0.1°. These are values found within the field area (Figure B.3). Therefore, the effects caused by the During the 20 year span, the declination was characterized by a RMS difference

were selected and are given in Table B.1. the region into 4 corners and the central region. Inclination and declination pairs forming the reduction-to-pole for the entire MA data set required the subsetting of modeling the RTP Therefore, the averages given in Figure B.3 were utilized as the basis for further process for this region. The approach that was selected for per-

of the change away from one corner to the next. Values within about 700 km  $(6.3^{\circ}N)$ Selection of five points to perform this adjustment was based upon the rapidity

Location	Latitude (°N)	) Longitude ( $^{\circ}$ W) Inclination ( $^{\circ}$ ) 1	Inclination (°)	Declination (°)
Southwest	Southwest 61.800	66.500	-38.81	81.16
Northwest	81.000	66.500	-78.36	86.22
Northeast	81.000	18.500	-31.18	83.45
Southeast	61.800	18.500	-19.63	74.19
Central	71.350	42.625	-43.47	81.13

km from these points represent a weighted average of the values generated by all the valid for approximately 700 km around each point. Values for points more than 700 reduction-to-pole is performed for each of these pairs, and the procedure is assumed lected at points inset from the four corners and at the middle of Figure B.3. Table B.1: Inclination and declination pairs for reduction-to-pole. below pairs. Values were se-

a blend of values obtained from different corner pairs. derived nearest each corner pair will be retained and the values in between must be range, the selected pair may not provide an adequate reduction-to-pole. So values x  $15.75^{\circ}E$ ) of each corner pair are assumed to be correctly reduced. Outside of that

more distant pair). derived from the closest pair is given more weight than the value derived from the accomplished by weighting the values according to proximity to their source grid the remaining grid values in the overlapped regins were linearly merged. This was given in Table B.1. The  $6.3^{\circ}N \times 15.75^{\circ}E$  patch around each pair was retained, and (i.e., if a point in question lies closer to one corner pair than another, then the value The 20 km upward-continued MA data were reduced-to-pole using the first 4 pairs

core region will enable an assessment of the errors involved in the linear merging Therefore, comparison of the 4-pair adjusted core region with middle pair adjusted These data should be adequately adjusted in terms of proximity to the central pair. was compared to values derived from the middle pair (last values in Table B.1). For purposes of comparison, the central portion of this 4-pair adjusted MA grid

4 corner pairs and would be the region having the highest probability of error. process. The values in the center of the 4-pair adjusted data are the furthest from all

25 to 100 nT [Verhoef et al., 1996]. data are well within the observation errors given with these data, which range from 14.0 nT). These results clearly indicate that the associated mismatches between these power ( $\sigma = 2.6 \text{ nT}$ ), little bias ( $\mu = 0.535 \text{ nT}$ ), and small amplitudes (max/min=9.6/-Differencing these cores generates a residual field that is characterized by low

and inclination the highest. regions are linearly merged, weighting the values derived from the closest declination around each of the five points given in Table B.1) and the data between these The final 5 pair adjusted grid retains the values within the range of  $6.3^{\circ}N \ge 15.75^{\circ}E$ As a final step, the middle pair derived grid was also added to the overall grid.

the gravity data. the magnetic anomaly source, a necessary step if these data are to be compared with of the Earth's magnetic field. of varying observations, and effectively reduced-to-pole to reduce the dipolar effects to remove biases and tilts, upward-continued to 20 km elevation to reduce the effects These MA data (Figure B.4) have now been hi-pass filtered at 400 km wavelengths The last aids in clarifying the location and nature of

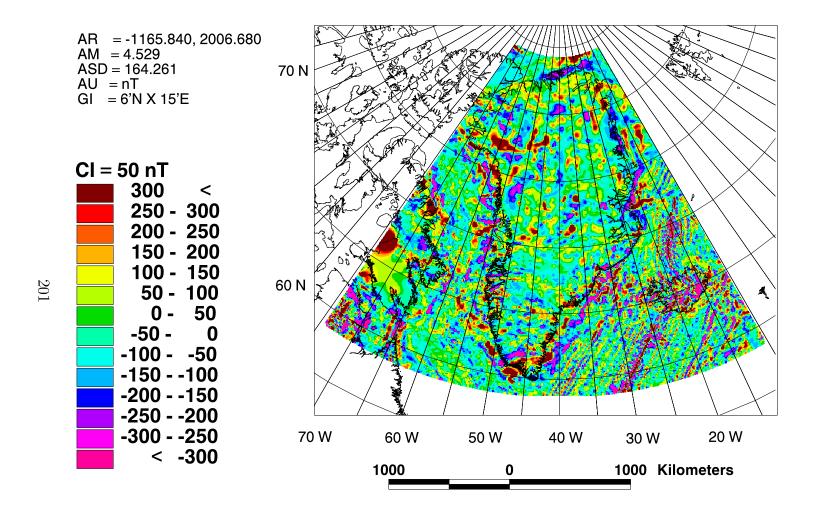


Figure B.1: Magnetic anomaly data provided by GSC from Open File 3125b for Greenland shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W.

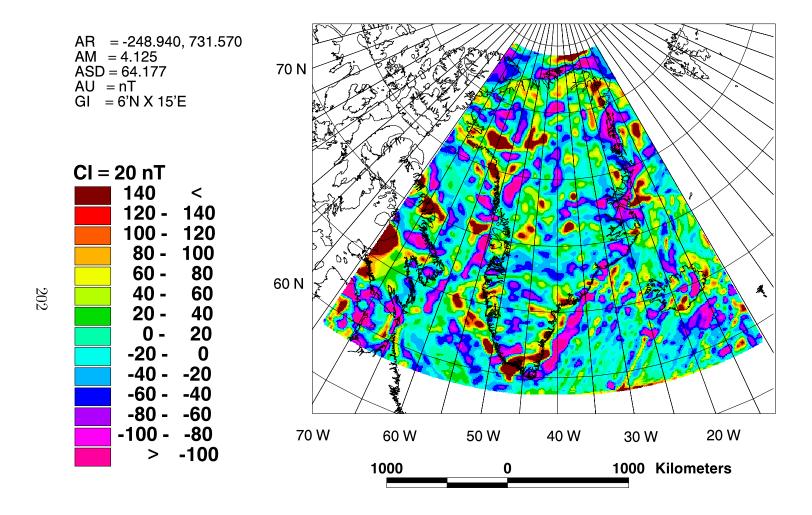


Figure B.2: Corrected magnetic anomaly data for Greenland shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W. Data are reduced-to-pole to clarify the location of magnetic sources and are upward continued to 20 km elevation for comparison with gravity data.

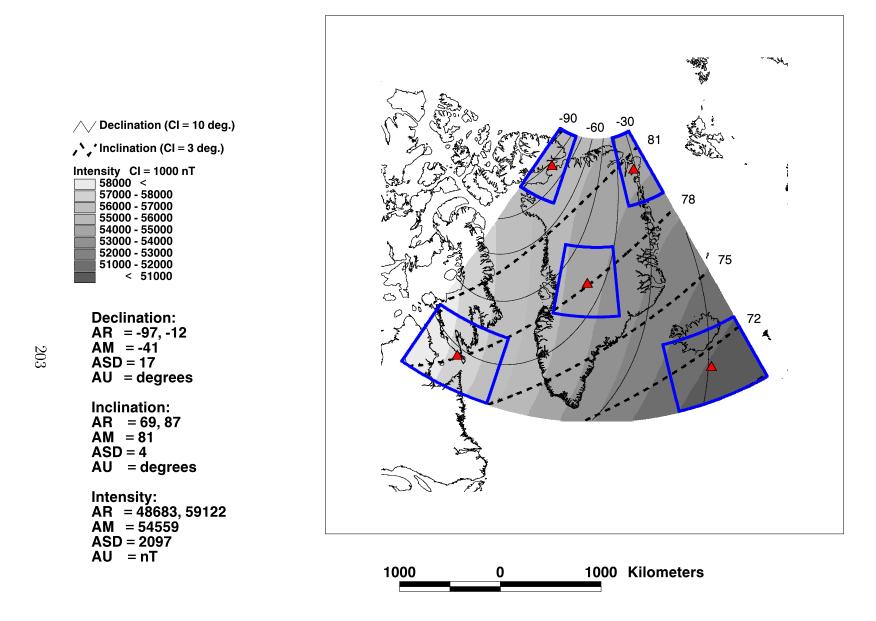


Figure B.3: Average IGRF derived Declination (solid line), Inclination (dashed line), and Intensity (grey scale) shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W. Data were obtained and averaged from the 1 degree IGRF models for epochs 1970-1990. The graticule is left out for clarity.

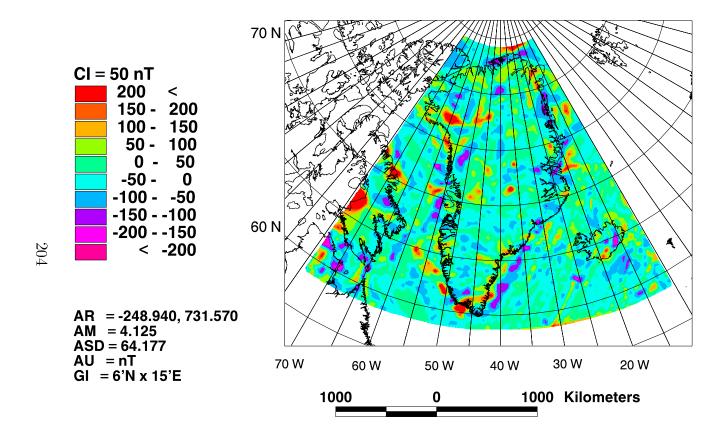


Figure B.4: MA reduced-to-pole and upward-continued to 20 km elevation shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W.

#### APPENDIX C

## Altimetry Enhanced Free-Air Gravity Anomalies

#### Abstract

for Geosat altimetry in this region. other available data for a 50% coherence at 35 km, which is the optimum expected show a coherence of 50% down to 50 km wavelength. They may be averaged with FAGA have been checked in the Gulf of Mexico against high density observations and the original FAGA to generate the final altimetry enhanced FAGA. These enhanced geoid surface is transformed into a residual FAGA surface, which is then merged with are merged to generate a final grid that has minimal trackline noise. This residual inversely transformed, averaged, and gridded. The two grids derived from each mode are selected. These wavelengths, which should reflect common geological sources, are wave domain, and those wavelengths with the highest correlation and retained power arately. Track pairs in each mode are geographically ordered, transformed into the for the Barents Sea. Residuals for ascending and descending modes are processed septo a reference geoid surface generated by integrating the Andersen & Knudsen FAGA ment to a reference free-air gravity anomaly (FAGA) grid. Altimetry is residualized graphically ordered, and their common components extracted to generate an adjust-The shortest wavelengths of ERS-1 altimeter data (168 day missions) are geo-

### C.1 Introduction

same altimeter data sets. their own methodology that have resulted in different FAGA estimations using the Rapp and Yi, 1996; Rapp et al., 1991; Sandwell and Smith, 1997]. Each author has (FAGA) for all ice-free, ocean areas [Andersen and Knudsen, 1998; Kim, J.-H., 1995; Satellite altimetry has been used extensively to derive free-air gravity anomalies

during different missions (Geosat, ERS-1, Topex, etc.). ing various weighting schemes to account for varying quality between data collected all available data sets have been incorporated into many of these analyses employaltimeter measurements obtained to close to coastal areas or in shallower seas. This is based upon quality assessment of the altimetry under those conditions. Additionally, Some criteria that authors use may eliminate or greatly reduce the weighting for

geological sources profiles to determine the similar static components, which are assumed to derive from approach. Therefore, an approach is offered here that assesses geographically adjacent the FAGA signal. Primarily, the data are reduced based entirely upon a numeric These schemes often do not take into account the geologic sources that generate

the longer wavelengths. However, Geosat altimetry extends only to  $\pm 72^{\circ}$  latitude. estimate the shorter wavelengths of the Earth's gravity field and all available data for 168-day missions (8 km at the equator). Therefore, it is desirable to use Geosat to much better between track spacing (3-4 km at the equator) than the combined ERS1 FAGA signals at double the spacing). Geosat altimetry from the GM mission has a because they will generate the best between track resolution (i.e., they could contain To that end, profiles that have the closest spacing between track are desired,

Therefore, the combined ERS-1 168 day missions were used, because they extend to

then spectrally compared to adjacent pairs residual profiles are then filtered to remove track noise, ordered geographically, and accomplished by determining and removing a reference geoid undulation field. lengths may be processed to determine a residual FAGA field. This is practically These data must have the longer wavelengths removed, so that the shorter wave-

mining the high frequency component of the Earth's gravity field. altimetry should register on both tracks. Geologic sources that have a signal larger than the between track spacing of the Retention of this signal is critical to deter-

remove-and-restore technique mentioned in Bašić and Rapp [1992] pilation of multiple altimetry sources. because subtle features are restored that were masked or removed during the com-The original reference FAGA are enhanced by the addition of this residual field, This added-value approach builds upon the

#### C.2Altimeter-Implied Free-Air Gravity Anomalies

good proxy indicator of the geoid undulation surface. (FAGA) from altimeter data. Both rely upon the fact the ocean surface is a relatively There are two primary approaches to the generation of Free-Air gravity anomalies

track derivative is that the long wavelength orbit errors are de-emphasized. to the desired FAGA [Sandwell and Smith, 1997]. The advantage to taking the along section points to estimate the deflection of the vertical, which can be directly related The first involves taking the along track derivative and interpolating this to inter-

problem of orbit errors is reduced in the following two ways. The residual profiles are undulation and then determine the FAGA from the geoid surface. approach [Kim, 1993] uses the altimetry to directly calculate The inherent ත geoid

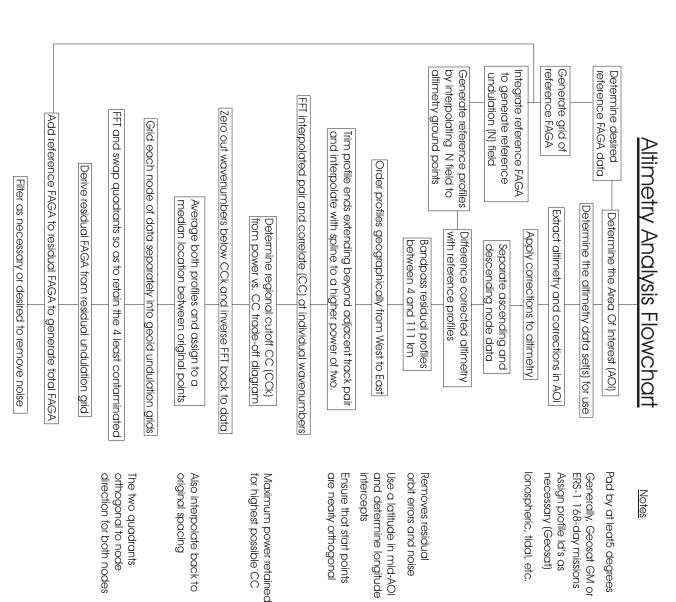
two grids reflect residual orbit errors, as well as track noise. repeated for all pairs within each track mode (ascending and descending). The final geographically ordered and compared to extract the most correlative static elements, which are assumed to have common crustal sources [Kim, 1993]. This process

plot to as a corrugated [Verhoef et al., 1997] or washboard [Kim et al., 1998] effect. generate a merged geoid undulation that has minimal errors, which have been referred is possible to retain the two least noisy quadrants from each grid's spectral amplitude minimal noise for the grids derived from different modes will be opposite. Therefore, it grid, two quadrants will have minimal noise. Due to orthogonality, the quadrants with The second means of reducing track noise is discussed in Kim et al., [1998]. In each and combine them into one spectrum. This may be inversely transformed to

that long wavelength errors are reduced first, then minimized further by taking the essentially vertical derivative The resulting geoid undulation grid is then used to generate the desired FAGA, by taking the first vertical derivative. The advantage to this approach is

generates a valid solution previously shown [Rapp et al., 1991] that use of this remove and restore technique of residual FAGA data that were derived from the residual altimetry. values from the altimetry and later restored the OSU91A FAGA to the Additionally, in his approach, Kim [1993] removed the OSU91A geoid undulation It has been final grid

(down to 20 km) than those from 360 degree coefficient models. ence geoid and FAGA data are utilized, which have information at shorter wavelengths and are outlined in the flowchart in Figure C.1. The principle difference is that refer-The enhancement techniques discussed next are a variant on the second approach



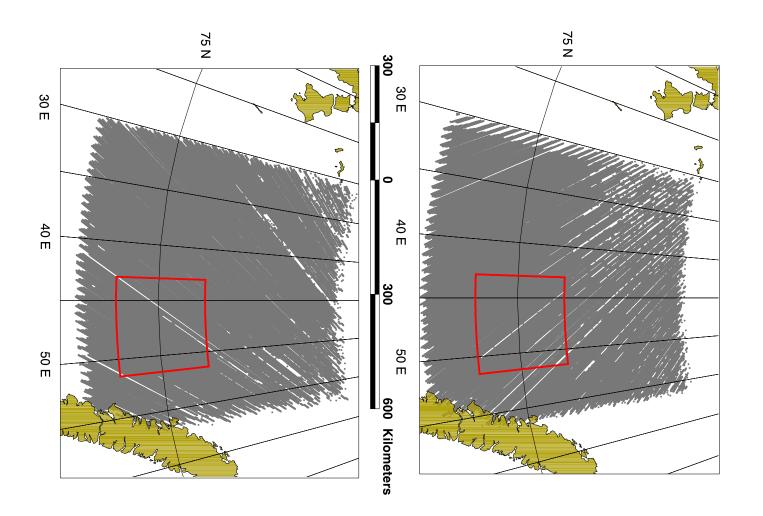
FAGA data is combined with the short wavelength data from the altimetry. generate enhanced FAGA. Ultimately, the long wavelength component of the reference Figure C.1: Altimetry analysis flowchart. Various processing steps are necessary to

## 7.3 A case study: The Barents Sea

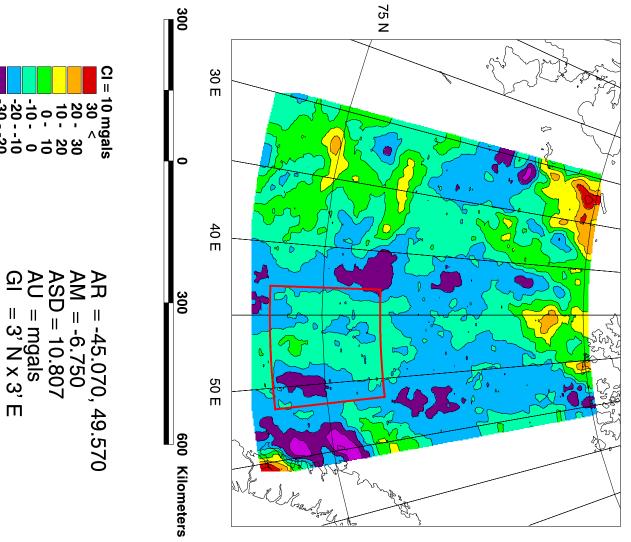
wavelength features of the FAGA. shelf sea the high frequency component of the FAGA field. Additionally, it is also a continental selected because of sufficient ERS1 altimetry coverage (Figure C.2) for determining  $\triangleright$ test area for this approach was selected in the Barents Sea. with shallow bathymetry that is likely to be well represented in the short This region was

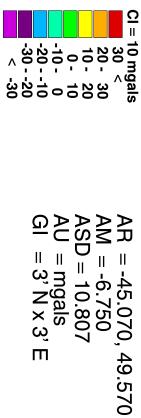
al., 1995]. be reliable down to a range of 20 - 50 km wavelengths depending on location [Yale et KAFAGA was derived from multiple altimetry data sets, which have ence for FAGA data for the shorter wavelengths that are of primary interest here. data and other data sets [Smith, 1997], however, they do provide an excellent refer-These techniques were implemented with altimetry from the ERS-1 168-day misand a reference FAGA grid from Andersen and Knudsen [1998] (KAFAGA), is shown in Figure C.3. There are some bias/tilt differences between KAFAGA been shown to

spikes noise and real signal. When the long wavelength differences and spikes were removed in Figure C.5 shows that the differences are primarily long wavelength, to residual orbit errors and tilt errors in the reference model. The bottom diagram long wavelength discrepencies between these two profiles, which may be attributed profile (solid) over the same groundpoints, and it can clearly be seen that there in Figure C.5 shows a reference undulation profile (dashed) and an observed altimeter undulation grid to the 1/10 second altimeter groundpoint locations. The top diagram Survey. KAFAGA was used to generate a reference geoid undulation grid (Figure C.4) and very high frequency features.  $\circ f$ Reference undulation profiles were generated by interpolating the reference ಶ 2D FFT transform provided by The Dru Smith of the National Geodetic latter may contain both observation along with



descending (bottom) altimeter coverage can be seen to be sufficient for the test area 45 E at sea level. (red box) and is shown in a Lambert Equal-Area Azimuthal Projection centered on Figure C.2: 168-day ERS-1 altimetry for the Barents Sea. Ascending (top) and





shown with the red box. Figure C.3: Andersen and Knudsen FAGA for the Equal-Area Azimuthal Projection centered on 45 Andersen and Knudsen FAGA for the Barents Sea shown in a Lambert E at sea level. The test area is

correlative (CC=0.999) and had little RMS difference (0.036 m). by filtering, the filter altimetry profile and the reference undulation profile were very

along track resolution, the minimal resolution from these data is about 4 km. track resolution while the 1/10 second sample interval (about 660 m) controls the km or greater for this region. Since the distance between tracks controls the crossassumption will have a negligible impact. Additionally, track spacing is generally 2 with the sea bottom. Therefore, any long wavelength errors that may result from this frequency field that will better represent the short wavelength features associated as orbit errors and filtered out. The intent here is to provide a more refined high For this analysis, all differences greater than 111 km wavelength will be treated

the fundamental equation of geodesy (Equation C.1) monic coefficients, whereas the models here are based directly upon observed data. sented by Bašić and Rapp [1992]. They removed fields derived from spherical har-The relationship of the geoid undulation field and the FAGA field are given through The removal and restoration of a reference field closely follows the technique pre-

$$\Delta g = -\frac{\partial T}{\partial r} - \frac{2T}{R} \tag{C.1}$$

where:  $\Delta g$  = free air gravity anomalies T = disturbing potential

radial direction

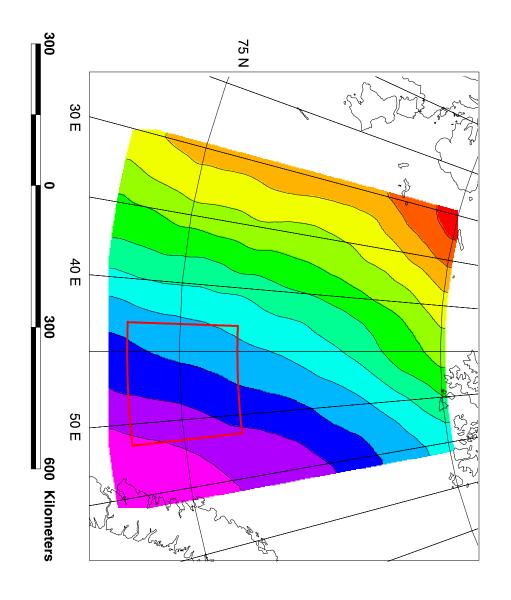
mean Earth radius =6371 km

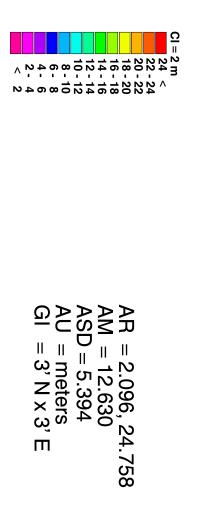
and Bruns' Formula Equation C.2) [Heiskanen and Moritz, 1967]

$$T = N\gamma \tag{C.2}$$

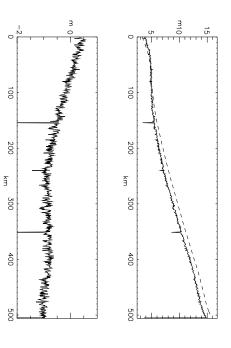
where: N = geoid undulation  $\gamma = normal gravity$ normal gravity

standard vertical derivative operator may be applied to the Fourier transform of the planar assumption is made based on the removal of the reference geoid, and a





level. The test area is shown with the red box. Figure C.4: Andersen and Knudsen FAGA-derived geoid undulations for the Barents Sea shown in a Lambert Equal-Area Azimuthal Projection centered on 45 E at sea

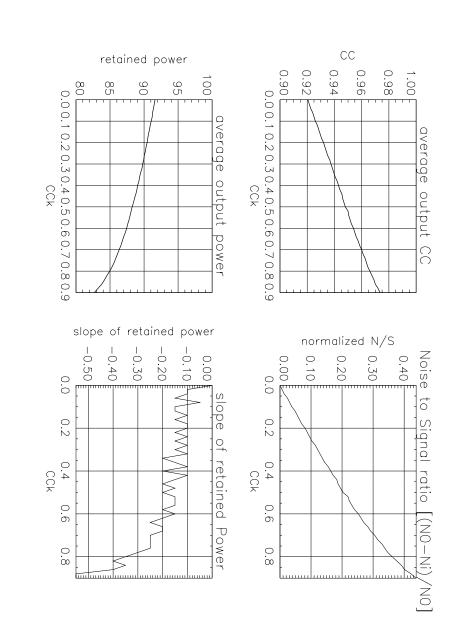


track 26138 and is sampled roughly every 660 meters. represent both noise and crustal signals. due to long wavelength (111+ km) errors, spikes, and higher frequency features that profiles (top) and the difference between them (bottom). Differences are primarily Figure C.5: Comparison of altimeter (solid) and reference geoid (dashed) derived Profile shown is ERS-1 168-day mission

added to the first to generate the desired FAGA. first component in Equation C.1. The second term may be easily determined and

of comparison. middle of the study area), thereby placing profiles in order from West to East for ease are geographically ordered by intercept longitude at a chosen latitude (usually in the until a final residual geoid undulation map is to be generated. The residual profiles parallel nature of the tracks. Analysis for profiles in both modes proceeds separately are separated by mode (ascending or descending), to take advantage of the subremove long wavelength errors and high frequency noise. As a first grouping, profiles profiles, which were then bandpassed filtered between 4 and 111 km wavelengths to Hence, reference profiles were removed from observed profiles to generate residual original two profiles squares estimate of the common signals and assigned to the median locations of the formed to reconstruct the track pair. The two profiles are averaged to produce a least occurs independently depending on whether they correlate better or worse than the those wavenumbers that correlate the least. The elimination of these wavenumbers very steep. This cutoff CC (CCk) is used to pass maximum power while removing established where the power drop off of the passed wavenumber components becomes gressively higher correlations is plotted (Figure C.6) and a critical cutoff point is effect on the overall power of selectively removing those wavenumbers that have prodifference gives the coefficient of correlation (CC) of each wavenumber pair. and which may be derived from the same geologic features. Small phase differences imply wavenumber components that are strongly correlated tween each pair of neighboring tracks are correlated based upon their phase difference. formed into the wavenumber domain. The individual wavenumber components Adjacent track pairs are then interpolated into a power of two and Fourier trans-CC(notch filtering). The remaining wavenumbers are then inversely trans-The cosine of the phase

justment. This method helps to reduce the corrugated effect associated with residual right and lower-left quadrants in the data domain). By retaining the two relatively the amplitude spectrum for descending data, which would pass through the uppertrack directions are apparent as a band in the quadrants orthogonal to the mode  $\operatorname{track}$ swapping method [Kim, 1993; Kim et al., 1998] is utilized instead of a crossover adfiles for an ascending or descending direction. To remove directional effects a quadrant versely transformed to generate Fourier spectra, where the effects of the individual These averaged profiles then most nearly represent residual geoid undulation proline bias when gridding. (eg.,ಶಾ band would appear in the upper-left and lower-right quadrants The two separate data modes are gridded and in-



CCk). inflection point is at about a CC of 0.8 to 0.9. For this reason, 0.9 was chosen for the of the power drop off (bottom-right diagram) more clearly shows that the maximal power is nearly linear except at the highest values (bottom-left diagram). The slope noise ratio are nearly linear (upper-right and upper-left diagrams) and the drop off in selection. Increase in overall track pair correlations and improvement in signal to Figure C.6: Power spectrum and trade-off diagram showing cutoff correlation (CCk)

final residual geoid undulation grid (Figure C.7). be generated. uncontaminated quadrants from each data mode, a composite Fourier spectrum may This composite spectrum is then inversely transformed to generate a

geographic latitude of the grid points. in a Fourier transform [Kim, 1993; Schwarz et al., 1990]. The data represented in Equation C.1. The second term is easily determined in the data domain knowing the in the wave (Figure C.7) are again Fourier transformed, and the first vertical derivative is taken (Figure C.8) through application of the fundamental equation of geodesy as expressed The residual geoid undulation may now be transformed into a residual FAGA grid domain. The result is inversely transformed to yield the first term in

select merging at this would not create obvious discontinuities in the combined spectra comparison of the radial power spectra of the residual FAGA and KAFAGA. Because is necessary to prevent the overlapped wavelengths from having too much power. A both radial power spectra contained about the same energy at the 55 km wavelength, value tween the 4 and 111 km wavelengths. To generate final predictions, it is necessary to This final residual FAGA grid represents that component of the FAGA field be- $\circ$ f a wavelength for merging these data with the reference KAFAGA grid. 55 km was selected as the most optimal merging wavelength based upon This

Figure C.9 with those in the red box portion of Figure C.3). but they can be seen to differ in the shorter wavelengths (compare the features in then added together to generate the enhanced FAGA grid (Figure C.9). This final grid (noise) that was being generated at those wavelengths. The two filtered grids were km in the residual FAGA were removed based upon the amount of incoherent signal FAGA were bandpass filtered between 16 and 55 km. Wavelengths between 4 and 16 strongly similar to KAFAGA (CC=0.93), because of the long wavelength features Therefore, KAFAGA were lowpass filtered KAFAGA at 55 km, and the residual

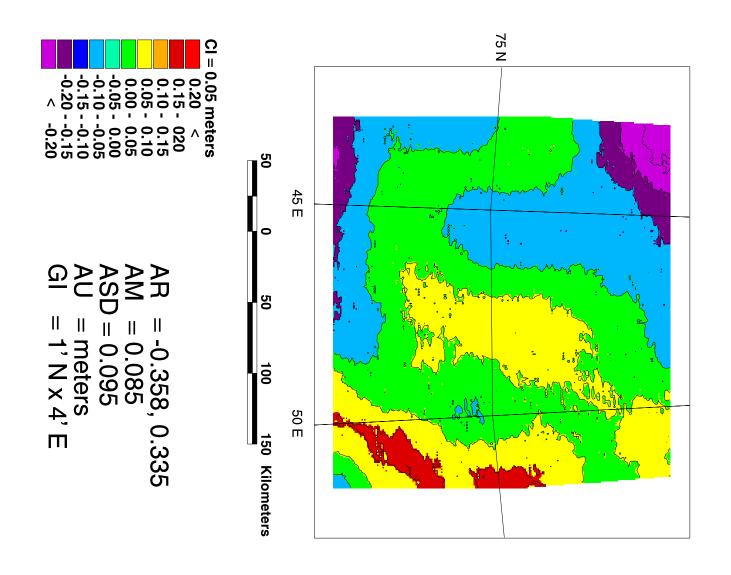
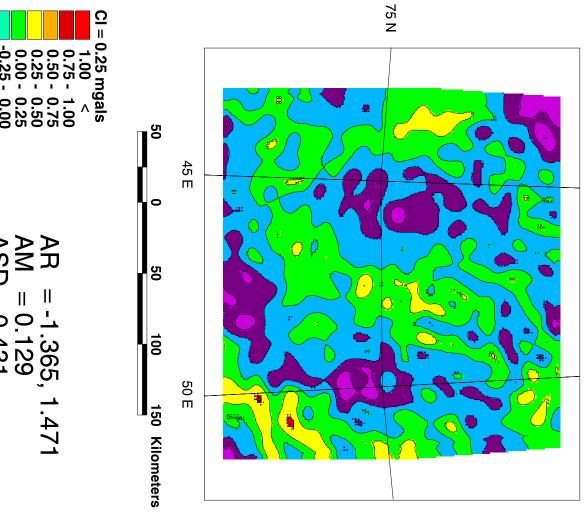


Figure C.7: Residual geoid undulation for Barents Sea test area shown in a Lambert Equal-Area Azimuthal Projection centered on 47 E at sea level.



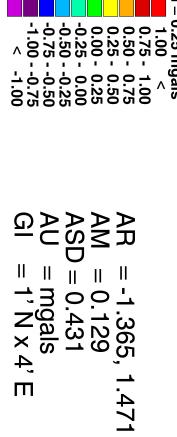
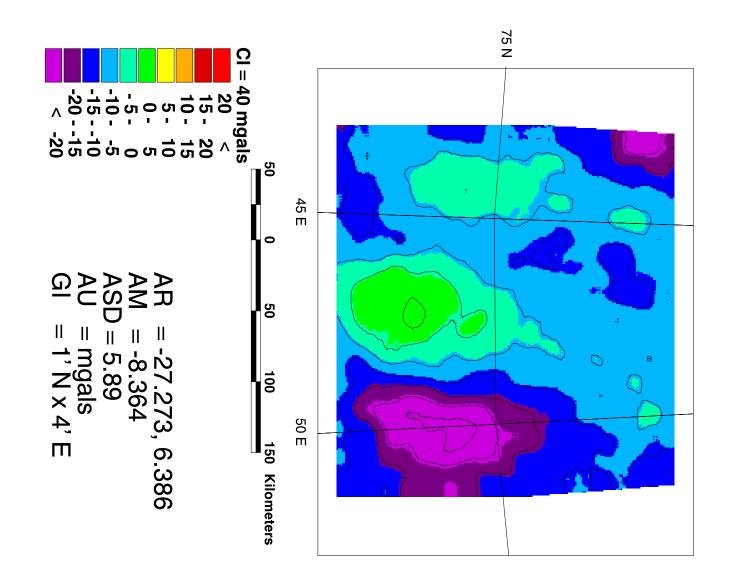


Figure C.8: Residual FAGA for Barents Sea test area shown in a Lambert Equal-Area Azimuthal Projection centered on 47 E at sea level.



to those in the red box in Figure C.3 show that the major features are still present. Area Azimuthal Projection centered on 47 E at sea level. Comparison of these features Figure C.9: Total enchanced FAGA for the Barents Sea shown in a Lambert Equal-

groundtrack coverage. optimum expected for Geosat data in this region [Yale et al., 1995] based upon the in processing and lowered the 50% coherence down to the 35 km, which is about the and Smith [1997] show agreement to the same level. Averaging the enhanced FAGA this reference FAGA down to the 50 km wavelength. FAGA data derived by Sandwell Gulf of Mexico region. about a 2 km resolution was available for comparison to the enhanced FAGA for the reference FAGA data set derived from bottom meter and shipborne observations at of southern Africa and in the South Atlantic [Roman and von Frese, 1998a]. with the Sandwell and Smith FAGA diminished irregularities derived from differences Efficacy of this procedure was tested in the Gulf of Mexico [Roman, 1996], offshore The enhanced FAGA data showed a coherence of 50% with

these two regions was comparable to that found for the Gulf of Mexico region. the agreement between the enhanced FAGA and the Sandwell and Smith FAGA for ilar reference FAGA data sets were not available for the other two regions, however, This would seem to substantiate this approach for the Gulf of Mexico region.

### C.4 Summary

from common geologic sources in both altimeter profiles. of geographically adjacent altimeter profiles. These components are assumed to derive data, which should be adequately described by all the combined data. The enhanced reference FAGA data set. This retains the long wavelength character of the reference The intermediate and longer wavelengths (55+km) remain the same as in the selected altimetry to refine the shortest wavelengths (10-55 km) of the Earth's gravity field. wasdiscussed. technique for enhancing available free-air gravity anomaly (FAGA) data sets wavelengths are derived by extracting the most correlative static components The enhancements retain the shortest wavelengths of the Geosat These value-added short

generate enhanced FAGA. wavelengths are then added back to the longer wavelengths of the reference FAGA to

limit of quality based upon Geosat's geographic coverage. with control data in the Gulf of Mexico - to a level estimated to be the theoretical data. This is evidenced by the further improvement in the merged data's correlation and Sandwell and Smith FAGA minimizes errors in both approaches and enhances the Smith [1997], which have been low pass filtered at 20 km. Merging of the enhanced The resulting enhanced FAGA data are comparable to those of Sandwell and

#### APPENDIX D

# Tabular control data for Greenland

the 3D models. The following data in Tables D.1 and D.2 were used to estimate the accuracies of

,	)								
•	1791.6	25.0	-999.0	1828.6	982027.5	313.14642	69.27756	cd-64	95027
	1856.4	32.9	-999.0	1893.4	982030.0	313.10162	69.52214	cd-38	95026
	1872.7	17.6	-999.0	1909.8	982019.2	313.06625	69.68202	cd-21	95025
	2081.7	37.2	-999.0	2119.5	982011.3	313.10965	70.31861	wp32	95029
	2075.2	38.5	-999.0	2112.5	982028.7	312.77301	70.56281	wp31	95030
	2060.4	32.5	-999.0	2097.4	982041.2	312.42819	70.80669	wp30	95031
	2062.2	44.9	-999.0	2098.9	982066.7	312.07367	71.04875	wp29	95032
	2065.9	51.9	-999.0	2102.2	982086.1	311.71234	71.29095	wp28	95033
	2074.9	49.1	-999.0	2111.0	982093.9	311.34229	71.53267	wp27	95034
	1993.5	50.8	-999.0	2029.7	982133.6	310.96066	71.77248	wp26	95035
	1987.4	47.8	-999.0	2023.8	982145.4	310.57001	72.01314	wp25	95036
	2187.9	79.9	-999.0	2224.3	982128.3	310.16968	72.25228	wp24	95037
	2143.1	59.4	-999.0	2178.9	982134.3	309.75842	72.49008	wp23	95038
	2040.9	58.6	-999.0	2075.6	982177.7	309.33380	72.72717	wp22	95039
	1941.3	47.6	-999.0	1974.6	982209.9	308.89990	72.96333	wp21	95040
	1849.0	39.5	-999.0	1881.0	982242.6	308.45438	73.19886	wp20	95041
	1797.7	25.4	-999.0	1828.9	982256.4	307.99582	73.43300	wp19	95042
	1807.8	36.0	-999.0	1838.8	982273.8	307.91098	73.63445	wp18	95043
	1828.7	42.2	-999.0	1859.5	982286.7	307.79199	73.90125	wp17	95044
	1858.0	37.1	-999.0	1888.6	982285.4	307.67041	74.16847	wp16	95046
	1866.6	44.6	-999.0	1897.1	982302.8	307.54504	74.43411	wp15	95047
	1858.7	27.2	-999.0	1888.9	982300.3	307.42026	74.70022	wp14	95048
	1921.7	38.0	-999.0	1951.4	982306.4	307.24557	75.01611	wp13	95053
	1957.6	30.4	-999.0	1987.0	982299.4	307.04007	75.27437	wp12	95052
	1900.7	38.8	-999.0	1929.8	982335.2	306.52286	75.49508	wp10	95051
	1917.0	47.9	-999.0	1945.9	982351.4	305.91260	75.77031	wp09	95050
	1876.2	48.2	-999.0	1904.1	982372.7	305.12424	75.95753	wp08	95054
	1755.8	26.2	-999.0	1782.5	982396.1	304.31516	76.14239	wp07	95049
	1642.8	1.0	-999.0	1668.5	982413.6	303.48199	76.32359	$^{ m wp06}$	96120
	1612.3	2.7	-999.0	1637.5	982432.4	302.63733	76.50394	wp05	96119
	1652.3	20.5	-999.0	1677.2	982445.1	301.75748	76.67939	wp04	95059
	1732.3	26.4	-999.0	1757.0	982434.9	300.64227	76.88794	wp03	95058
	1829.9	32.9	-999.0	1854.7	982416.6	299.93195	77.02370	wp02	95056
	1873.0	45.7	-999.0	1897.8	982420.6	298.96100	77.13975	$\operatorname{gits}$	95057
	1882.0	45.7	-999.0	1906.8	982419.5	298.87903	77.18044	wp01	95055
	(m)	(mgals)	(m)	(m)	$(\overline{\mathrm{mgals}})$	(deg. E)	(deg. N)	name	num.
	elev.	$\Delta g$	thick.	elev.	gravity	longitude	latitude	station	id.
•	orth.	FAGA	ice	ell.	absolute	location	loca	point	pc

normal gravity value free-air corrected to the observation elevation. A value of -999.0 in the radar ice thickness column indicates a null value for that location. Table D.1: BPRC-derived location, gravity and ice thickness data in Greenland. Free-air gravity anomalies were derived by differencing the absolute gravity value with a

Table D.1: Continued

1010	t	logi	1.	, h = , h, t ,	<u>,</u> =	}	V U V	ا الم
id.	Station	latitude	e longitude	gravity	elev.	thick.	$\Delta a$	elev.
num.	name	(deg. N)	$(\deg. E)$	(mgals)	(m)	(m)	(mgals)	(m)
95028	cd-107	68.87097		982013.3	1808.8	-999.0	29.3	1771.4
95019	c136	68.59683	313.24966	981971.8	1826.1	-999.0	10.0	1788.2
95020	c159	68.37881	313.29138	981945.6	1873.1	-999.0	11.8	1834.8
95021	c182	68.16103	313.35971	981944.7	1875.7	-999.0	25.2	1836.7
95022	c206	67.93581	313.43048	981923.4	1923.2	-999.0	32.8	1883.2
95023	c231	67.70025	313.48917	981913.3	1955.8	-999.0	47.8	1914.8
95024	c250	67.51965	313.49341	981906.1	1972.1	-999.0	57.1	1930.4
96001	f2	67.43945	313.31592	981913.6	1936.0	-999.0	58.6	1894.4
96002	<b>3</b>	67.18000	313.13916	981911.0	1914.0	-999.0	66.1	1872.0
96003	f4	66.91888	312.97006	981900.4	1882.0	-999.0	62.6	1839.9
96004	f5	66.65833	312.80334	981879.8	1898.0	-999.0	64.1	1856.1
96005	f6	66.39735	312.63971	981847.8	1922.8	-999.0	57.0	1881.3
96006	f7	66.13618	312.47839	981819.8	1932.1	-999.0	49.4	1890.8
96007	f8	65.87498	312.32336	981792.0	1927.6	-999.0	37.8	1886.6
96008	f9	65.61368	312.16986	981780.9	1919.9	-999.0	42.1	1879.2
96009	f10	65.35204	312.01910	981769.1	1943.9	-999.0	55.6	1903.4
96010	f11	65.08971	311.87103	981739.6	1944.8	-999.0	44.6	1904.3
96011	f12	64.82775	311.72742	981752.6	1888.3	-999.0	58.4	1848.0
96012	f13	64.56610	311.58554	981762.9	1831.3	-999.0	69.3	1791.6
96013	f14	64.29765	311.56506	981746.1	1784.4	-999.0	57.1	1745.0
96014	f15	64.02806	311.54568	981675.8	1876.4	-999.0	34.4	1837.3
96015	f16	63.75964	311.52640	981669.0	1903.0	-999.0	55.0	1863.9
96016	f17	63.49043	311.50735	981670.3	1894.9	-999.0	73.2	1855.6
96017	f18	63.22149	311.48941	981574.9	1912.0	-999.0	2.7	1872.3
96018	f19	62.95289	311.47156	981660.1	1911.8	-999.0	107.6	1871.9
96019	f20	62.72915	311.80081	981626.9	1952.0	-999.0	103.2	1911.0
96020	f21	62.50461	312.12436	981594.2	1855.2	-999.0	57.2	1814.0
96021	f22	62.28048	312.43985	981579.4	1832.4	-999.0	52.1	1791.3
96022	f23	62.05419	312.75473	981645.9	1900.5	-999.0	156.6	1858.9
96023	f24	61.82769	313.06296	981618.1	1954.2	-999.0	162.5	1911.9
96024	f25	61.81112	313.63046	981619.8	2123.0	-999.0	217.6	2078.8
96025	f26	61.79196	314.19711	981581.9	2220.8	-999.0	211.2	2175.2
96027	f28	61.74627	315.32690	981588.0	2184.8	-999.0	209.7	2136.4
96028	f29	61.72227	315.89523	981597.5	2091.6	-999.0	192.2	2042.5
96029	f30	61.93361	316.24533	981617.2	2172.7	-999.0	221.0	2123.2
96030	f31	62.20261	316.25781	981614.1	2160.5	-999.0	194.0	2110.5
96031	f32	62.47179	316.27090	981560.2	2122.9	-999.0	108.3	2072.9
96032	f33	62.74995	315.86636	981612.3	2465.2	-999.0	245.4	2414.2
96033	f34	63.03033	316.12515	981596.3	2392.8	800.0	186.5	2341.6
96034	f35	63.24503	316.32718	981599.2	2535.7	1005.0	217.8	2484.2

Table D.1: Continued

po	point	loca	location	absolute	ell.	ice	FAGA	orth.
1d.	station	latitude	longitude	gravity	elev.	thick.	$\Delta g$	elev.
num.	name	(deg. N)	(deg. E)	(mgals)	(m)	(m)	(mgals)	(m)
96035	f36	63.49422	316.56540	981609.0	2563.0	1110.0	217.9	2511.2
96036	f37	63.73913	316.80472	981616.6	2628.6	1206.0	228.1	2576.9
96037	f38	63.98901	317.04977	981560.3	2477.2	1204.0	107.1	2426.0
96038	f39	64.23389	317.30005	981620.9	2359.8	987.0	113.9	2309.1
96039	f40	64.47779	317.55508	981677.1	2072.6	1099.0	64.3	2022.6
96040	f41	64.74725	317.58133	981650.7	2203.8	1591.0	59.4	2154.1
96041	f42	65.01582	317.60895	981712.5	2119.1	1367.0	76.3	2069.7
96042	f43	65.28552	317.63864	981730.8	2003.5	1628.0	40.2	1954.4
96043	f44	65.55267	317.66553	981745.7	2035.0	1667.0	46.6	1986.1
96044	f45	65.78459	318.00610	981763.9	2035.6	1622.0	49.1	1987.0
96045	f46	66.01218	318.34552	981754.0	2101.8	1824.0	44.2	2053.3
96046	f47	66.24221	318.70358	981756.8	2129.8	1762.0	40.2	2081.1
96047	f48	66.46911	319.05737	981765.9	2209.4	-999.0	58.8	2160.3
96048	f49	66.69603	319.41699	981771.7	2194.7	-999.0	45.0	2145.3
96049	f50	66.92274	319.78754	981772.8	2186.7	-999.0	28.8	2137.2
96050	f51	67.14787	320.16406	981764.6	2269.7	-999.0	31.5	2220.0
96051	f52	67.31219	320.71738	981747.0	2348.9	-999.0	27.7	2298.5
96052	f53	67.47318	321.27255	981709.5	2482.6	-999.0	21.1	2431.0
96053	f54	67.63355	321.83942	981706.8	2628.9	-999.0	53.2	2576.0
96054	f55	67.79007	322.41507	981668.2	2754.9	-999.0	43.6	2700.7
96055	f56	67.94671	322.99634	981656.0	2886.4	2096.0	62.0	2831.2
96056	f57	68.09963	323.58398	981665.0	2977.9	1654.0	89.6	2921.8
96057	f58	68.25179	324.18500	981701.0	2931.8	1406.0	101.8	2875.2
96058	f59	68.53062	324.25504	981691.3	2799.4	2300.0	33.9	2743.4
96059	f60	68.81191	324.33093	981761.6	2753.9	1729.0	72.9	2698.4
96060	f61	69.09526	324.41483	981780.4	2696.2	1900.0	56.6	2640.9
96061	f62	69.37518	324.49200	981820.9	2720.5	1989.0	87.8	2665.1
96062	f63	69.52222	325.17978	981846.6	2720.3	1463.0	104.7	2664.2
96063	f64	69.66646	325.87146	981862.5	2664.4	1335.0	94.8	2608.2
96064	f65	69.80849	326.57544	981820.7	2696.0	1700.0	54.3	2639.5
96065	f66	69.94885	327.29025	981812.6	2769.4	1800.0	60.6	2712.3
96066	f67	70.22163	327.38062	981798.6	2871.5	2077.0	62.2	2814.7
96067	f68	70.49577	327.47382	981828.1	2870.0	1914.0	75.6	2813.9
96068	f69	70.76944	327.56656	981845.7	2816.3	2001.0	61.0	2761.3
96069	f70	71.04166	327.66376	981859.6	2785.3	2205.0	50.1	2731.3
96070	f71	71.31603	327.76578	981897.3	2747.7	2092.0	60.9	2694.3
96071	f72	71.58830	327.86819	981912.3	2701.8	2153.0	46.8	2649.0
96073	f75	72.13501	328.08469	981891.5	2749.5	2416.0	11.4	2698.2
96072	f74	72.40752	328.19888	981948.3	2855.6	1500.0	86.6	2804.9
96074	f76	72.67838	328.32465	981915.9	2903.6	1615.0	54.9	2853.2

po	point	loca	location	absolute	ell.	ice	FAGA	orth.
id.	station	latitude	longitude	gravity	elev.	thick.	$\Delta g$	elev.
num.	name	(deg. N)	(deg. E)	(mgals)	(m)	(m)	(mgals)	(m)
96075	f77	72.95300	328.43372	981932.0	2912.2	1747.0	59.5	2861.9
96076	f78	73.22584	328.55673	981947.5	2916.8	1348.0	62.7	2866.8
96077	f81	73.50395	328.69742	981981.2	2890.1	823.0	74.2	2840.7
96078	f82	73.77976	328.93805	982020.5	2821.1	686.0	78.7	2772.4
96079	f83	74.05439	329.18698	982006.9	2773.6	955.0	37.2	2725.5
96080	f84	74.32832	329.45142	982063.0	2729.3	1442.0	66.6	2681.8
96081	f85	74.60120	329.73035	982096.1	2655.5	1357.0	64.1	2608.7
96082	f86	74.87357	330.01639	982087.8	2663.3	1796.0	45.6	2617.0
96083	f87	75.14738	330.30682	982097.2	2585.5	2085.0	18.6	2539.3
96084	f88	75.41988	330.61261	982183.0	2536.9	1648.0	77.2	2490.1
96085	f89	75.69181	330.92831	982226.0	2462.8	2022.0	85.4	2415.3
96086	f90	75.96321	331.25735	982290.7	2365.0	1376.0	108.2	2317.2
96087	f91	76.23466	331.60245	982326.8	2287.8	1428.0	109.0	2240.4
96088	f92	76.50513	331.95523	982402.9	2178.0	831.0	139.9	2131.7
96089	f96	76.77868	332.32922	982412.7	2049.0	1530.0	98.8	2004.2
96090	f97	77.02991	331.89987	982442.3	1912.1	1457.0	76.1	1868.7
96091	f98	77.28230	331.49570	982456.6	1854.1	1705.0	62.6	1812.0
96092	f99	77.53101	331.02560	982484.3	1817.3	1440.0	69.2	1776.2
96093	f100	77.77999	330.55496	982498.9	1759.1	1523.0	56.4	1718.8
96094	f101	78.02887	330.06564	982522.4	1729.2	1387.0	61.4	1689.8
96095	f102	78.32616	329.44821	982504.8	1758.6	1674.0	42.1	1720.1
96096	f103	78.62091	328.80927	982517.8	1768.7	1565.0	47.7	1730.7
96097	f104	78.91445	328.13513	982522.5	1800.1	1585.0	51.9	1762.5
96098	f105	79.20844	327.41010	982518.4	1832.6	1673.0	47.9	1795.2
96099	f106	79.49748	326.66086	982518.5	1889.8	1766.0	56.1	1852.1
96100	f107	79.78625	325.86435	982531.9	1941.6	1769.0	76.2	1903.7
96101	f108	80.07324	325.02707	982528.7	1956.2	1792.0	68.5	1918.9
96102	f109	80.35745	324.13358	982521.7	1914.6	1631.0	40.1	1878.3
96103	f110	80.63997	323.18909	982544.4	1871.8	1719.0	41.1	1836.1
96104	f113	80.91976	322.17819	982597.2	1847.6	1322.0	78.4	1812.3
96105	f114	80.75153	320.08655	982516.2	2008.5	1749.0	51.9	1973.9
96106	f115	80.57153	318.06104	982494.7	2035.1	1711.0	43.9	2001.7
96107	f116	80.38071	316.10248	982485.8	2033.1	1709.0	40.1	2000.7
96108	f117	80.17941	314.23758	982485.1	2032.3	1823.0	45.2	2001.0
96109	f118	79.96825	312.44284	982466.2	2028.6	1963.0	31.6	1998.3
96110	f119	79.74799	310.72461	982443.7	2021.2	2376.0	13.8	1991.4
96111	f120	79.51905	309.07764	982475.4	1963.7	1860.0	35.1	1934.3
96112	f121	79.28233	307.50296	982466.1	1959.8	1967.0	32.4	1931.1
96113	f122	79.03761	305.99194	982452.0	1968.1	2035.0	29.1	1939.9
96114	f123	78.78490	304.55078	982448.0	1981.5	1945.0	37.9	1953.4
96115	fh01	78.25407	302.24460	982435.9	2028.8	1894.0	59.1	2000.3
96116	fh02	77.97932	301.37094	982439.2	2074.9	1709.0	86.8	2046.7
96117	fh03	77.70087	300.52982	982393.4	2082.1	1949.0	53.6	2055.0
	3	77 /1075	200 72/12	000001 0	2011 <i>a</i>	10000	<b>91</b> 0	

point		Ekholm		NIMA	A
id. no.	ice surface	ice bottom	ice thickness	ice surface	FAGA
	(m)	(m)	(m)	(m)	(mgals)
95055	1877.8	564.4	1313.4	1875.2	31.5
95057	1870.0	555.0	1315.0	1869.5	32.3
95056	1819.4	384.5	1434.9	1814.5	19.4
95058	1745.5	330.1	1415.4	1741.2	19.3
95059	1662.5	316.7	1345.8	1661.8	11.3
96119	1627.4	184.9	1442.5	1624.5	-12.9
96120	1686.0	305.5	1380.5	1675.6	-7.0
95049	1740.1	450.0	1290.1	1746.5	16.8
95054	1854.9	593.4	1261.5	1850.7	38.7
95050	1914.7	506.9	1407.8	1910.8	36.2
95051	1915.0	454.1	1460.9	1909.5	22.6
95052	1941.8	418.9	1522.9	1937.3	10.1
95053	1915.2	511.0	1404.2	1911.1	22.1
95048	1872.1	452.9	1419.2	1871.3	26.3
95047	1863.1	437.1	1426.0	1859.6	28.9
95046	1851.1	385.7	1465.4	1850.3	26.0
95044	1832.9	419.8	1413.1	1829.2	27.9
95043	1803.1	412.2	1390.9	1800.1	16.9
95042	1789.9	398.5	1391.4	1784.4	18.4
95041	1847.5	480.4	1367.1	1846.5	24.1
95040	1941.8	603.8	1338.0	1942.7	18.7
95039	2046.9	695.2	1351.7	2040.2	43.0
95038	2144.0	759.3	1384.7	2141.9	44.6
95037	2183.0	897.0	1286.0	2175.3	64.8
95036	1987.4	717.8	1269.6	1980.1	36.9
95035	1977.1	702.6	1274.5	1972.7	44.6
95034	2073.2	601.1	1472.1	2071.1	37.2
95033	2067.3	313.9	1753.4	2064.1	19.1
95032	2055.1	251.4	1803.7	2052.4	27.4
95031	2060.7	242.9	1817.8	2058.5	23.1
95030	2070.9	217.4	1853.5	2067.0	19.3
95029	2075.0	167.0	1908.0	2073.3	28.2
95025	1891.4	-4.9	1896.3	1901.7	12.5
95026	1855.0	16.5	1838.5	1851.1	12.4
95027	1794.7	30.3	1764.4	1805.4	20.6
95028	1769.3	22.4	1746.9	1768.0	9.9

Table D.2: Height, thickness and FAGA data interpolated from 3D data for the BPRC data points on Greenland.

Table D.2: Continued

i	10000	110.0	11000.1	EOH+++	00000
90.± 01 2	2449.3 2508 0	1975 0	1935.9	2511 1 2511 1	90035
06.4	2//0.2	1911 /	1967 5	2478 O	06037
132.2	2373.3	926.0	1460.2	2386.2	96033
94.1	2425.8	1027.7	1390.8	2418.5	96032
94.3	2134.4	634.8	1482.2	2117.0	96031
92.9	2151.1	910.9	1244.3	2155.2	96030
102.6	2101.6	804.7	1349.7	2154.4	96029
76.0	2189.7	718.7	1479.5	2198.2	96028
84.0	2142.9	715.6	1421.2	2136.8	96027
115.0	2090.7	434.0	1667.6	2101.6	96025
91.2	2060.3	361.4	1707.7	2069.1	96024
63.4	1898.4	268.9	1645.7	1914.6	96023
46.3	1865.5	438.0	1433.8	1871.8	96022
42.3	1799.3	525.3	1278.8	1804.1	96021
60.3	1807.6	535.3	1278.3	1813.6	96020
79.6	1887.3	471.7	1432.4	1904.1	96019
67.9	1841.2	285.7	1567.5	1853.2	96018
27.3	1894.4	281.6	1623.6	1905.2	96017
58.3	1866.9	423.7	1460.6	1884.3	96016
43.0	1842.0	696.2	1146.3	1842.5	96015
39.2	1790.4	956.0	854.2	1810.2	96014
37.2	1766.9	895.7	879.3	1775.0	96013
57.7	1793.1	842.9	949.2	1792.1	96012
55.8	1856.2	910.3	950.2	1860.5	96011
51.3	1899.8	1086.8	809.4	1896.2	96010
41.4	1890.1	1239.2	658.6	1897.8	96009
59.7	1871.6	1288.2	585.8	1874.0	96008
30.2	1873.8	1324.9	549.2	1874.1	96007
50.2	1885.3	1366.6	519.9	1886.5	96006
37.0	1892.8	1318.0	573.4	1891.4	96005
37.4	1844.6	1335.5	513.2	1848.7	96004
61.4	1836.5	1396.5	440.8	1837.3	96003
60.6	1874.7	1478.4	398.3	1876.7	96002
44.1	1894.0	1604.9	289.3	1894.2	96001
42.9	1927.3	1685.7	242.6	1928.3	95024
36.7	1907.1	1724.2	187.5	1911.7	95023
23.8	1879.0	1767.9	115.4	1883.3	95022
8.4	1832.3	1829.8	6.8	1836.6	95021
-6.4	1827.4	1945.9	-112.9	1833.0	95020
-6.0	1796.1	1939.0	-152.0	1787.0	95019
(mgals)	(m)	(m)	(m)	(m)	
FAGA	ice surface	ice thickness	ice bottom	ice surface	id. no.
A	NIMA		$\operatorname{Ekholm}$		point

Table D.2: Continued

point		Ekholm		NIMA	A
id. no.	ice surface	ice bottom	ice thickness	ice surface	FAGA
96036	$\frac{(2567.1)}{2567.1}$	1300.1	1267.0	2563.9	105.3
96037	2431.4	1250.5	1180.9	2439.1	81.0
96038	2304.6	1331.1	973.5	2311.1	96.4
96039	2020.2	1009.6	1010.6	2026.8	57.1
96040	2146.1	1042.1	1104.0	2148.5	52.5
96041	2086.3	956.5	1129.8	2084.8	68.2
96042	1952.2	533.6	1418.6	1953.3	64.0
96043	1981.7	368.0	1613.7	1981.1	39.6
96044	1988.6	341.6	1647.0	1991.6	27.3
96045	2042.3	346.1	1696.2	2042.6	25.5
96046	2090.9	371.9	1719.0	2091.5	22.7
96047	2153.2	343.5	1809.7	2152.0	41.5
96048	2146.4	186.3	1960.1	2148.9	20.8
96049	2146.2	157.4	1988.8	2146.1	17.7
96050	2227.5	231.3	1996.2	2233.4	17.0
96051	2306.8	298.8	2008.0	2307.8	13.0
96052	2429.7	385.7	2044.0	2425.0	17.4
96053	2572.4	526.6	2045.8	2571.7	34.6
96054	2719.3	780.2	1939.1	2718.5	51.7
96055	2829.7	965.8	1863.9	2824.1	65.2
96056	2924.6	1240.1	1684.5	2922.5	85.1
96057	2881.1	1371.6	1509.5	2874.2	82.4
96058	2736.6	895.1	1841.5	2742.8	57.9
96059	2708.1	867.7	1840.4	2707.8	54.3
96060	2646.1	980.7	1665.4	2650.0	61.9
96061	2657.1	876.4	1780.7	2660.1	82.8
96062	2666.1	967.0	1699.1	2659.7	103.6
96063	2601.5	944.6	1656.9	2610.9	69.5
96064	2646.2	890.0	1756.2	2645.0	43.0
96065	2711.1	947.7	1763.4	2706.8	54.6
96066	2811.7	946.2	1865.5	2812.3	51.9
96067	2813.2	924.2	1889.0	2814.7	53.8
96068	2762.4	745.1	2017.3	2763.9	34.1
96069	2733.5	649.7	2083.8	2733.1	29.4
96070	2689.5	670.3	2019.2	2692.4	38.7
96071	2646.2	653.0	1993.2	2646.7	35.1
96073	2800.6	868.0	1932.6	2802.2	19.8
96072	2732.3	697.5	2034.8	2736.2	20.0
96074	2851.0	1003.2	1847.8	2853.6	38.0
96075	2861.0	1137.8	1723.2	2861.0	29.2

able
D.2:
Continued

	בויף יוים		MIN	۸
		io +hialmoss		FAC A
		(m)		(mgals)
2847.5	1309.7	$1\overline{537.8}$	$28\overline{45.9}$	42.1
2824.2	1518.8	1305.4	2827.6	59.1
2767.8	1581.7	1186.1	2769.3	63.3
2725.0	1430.6	1294.4	2726.1	16.7
2691.4	1366.0	1325.4	2690.6	38.7
2618.7	1160.9	1457.8	2621.7	32.2
2598.2	917.3	1680.9	2599.1	28.5
2539.2	833.9	1705.3	2541.0	-4.8
2489.5	717.6	1771.9	2490.9	45.6
2416.2	737.3	1678.9	2415.4	85.3
2300.2	850.3	1449.9	2303.8	93.5
2241.1	839.1	1402.0	2239.3	91.1
2128.9	802.7	1326.2	2132.7	110.5
2001.1	699.7	1301.4	1999.9	85.7
1865.9	534.6	1331.3	1869.6	79.1
1816.5	407.7	1408.8	1824.0	50.0
1767.2	277.8	1489.4	1769.7	38.7
1719.6	137.1	1582.5	1718.4	40.5
1678.8	148.1	1530.7	1681.1	44.1
1707.0	226.6	1480.4	1709.8	30.4
1732.5	162.9	1569.6	1730.7	26.6
1764.0	230.4	1533.6	1765.4	36.0
1792.2	273.9	1518.3	1791.8	25.4
1855.6	211.1	1644.5	1855.9	49.2
1905.0	216.4	1688.6	1905.0	61.9
1920.0	179.3	1740.7	1920.0	53.1
1878.9	$\frac{173.5}{217.4}$	1705.4	1879.7	32. 32. 32. 32. 32. 32. 32. 32. 32. 32.
1004.0	517.3	1905.0	1000.4	و برو د. او
1073 3	550 B	1413.5	1073 6	58.4
2002.2	702.7	1299.5	2001.4	39.6
2001.7	672.8	1328.9	2002.1	29.3
2000.0	496.1	1503.9	1998.2	23.4
1999.0	342.5	1656.5	2001.5	18.3
1992.7	226.1	1766.6	1993.1	16.2
1942.7	161.9	1780.8	1942.6	26.2
1930.9	137.9	1793.0	1930.7	22.1
1941.2	113.1	1828.1	1941.6	19.7
1953.8	138.7	1815.1	1953.8	27.1
2004.5	115.8	1888.7	2003.7	49.1
2051.9	235.9	1816.0	2053.7	75.2
2059.2	229.5	1829.7	2058.3	51.0
1988.0	354.7	1633.3	1985.2	24.0
	ice surface (m) 2847.5 2847.5 2847.5 2847.5 22691.4 2618.7 2539.2 2489.5 2416.2 2530.2 2241.1 2128.9 2001.1 1865.9 1876.2 17767.0 1792.2 1878.9 1812.5 1764.0 1792.2 1834.0 1878.9 1812.5 1973.3 2002.2 2001.7 2193.9 1841.7 1942.7 1942.7 1942.7 1942.7 1953.8 2004.5 2059.2 1988.0	surface ice (m) 691.4 691.4 691.4 691.4 691.4 691.4 691.4 691.4 691.4 691.5 618.7 618.7 618.7 618.7 618.7 618.7 618.7 618.7 618.7 618.7 618.7 618.7 618.8 65.9 816.5 678.8 825.6 678.8 825.6 678.8 629.0 990.0 990.0 990.0 990.7 618.2 5 973.3 973.3 973.3 973.3 973.3 973.3 973.3 973.3 973.8 618.7 618.8 618.7 618.8 618.2 618.8 618.2 6	Ekholm surface ice bottom ice (m)	Ekholm           surface         ice bottom         ice thickness         io           (m)         (m)         (m)         (m)           824.2         1518.8         1305.4         137.8           824.2         1186.1         1186.1         1294.4           6618.7         1160.9         1457.8         1294.4           6618.7         1160.9         1457.8         1294.4           6618.7         1160.9         1457.8         1294.4           6618.7         1160.9         1457.8         1294.4           6618.7         1160.9         1457.8         1680.9           539.2         833.9         1705.3         1480.9           4416.2         737.3         1678.9         1449.9           241.1         839.1         1402.0         1128.9         1301.4           865.9         534.6         1331.3         1449.9           241.1         699.7         1301.4         1402.0           1128.9         1331.3         1301.4         1408.8           767.2         277.8         1489.4         1530.7           764.0         230.4         1533.6         1548.4           878.9

#### APPENDIX E

# Seismic Control Depths for Greenland

#### Abstract

may arise where erroneous density assumptions were used. Moho model may be reliablely used to model regional effects, although local problems and the percentile RMS difference is 13.1%. The overall agreement indicates that the the correlation is 0.95, the mean difference is 0.4 km, the RMS difference is  $\pm 3.4$  km, possible errors in the locations of the seismic estimates. If these 6 are excluded, then arose primarily due to deviations away from the uniform density assumptions and to seismic estimates, was 18.3% for all values. There were six problematic values that resulting percentile, representing the magnitude of the difference with respect to the made dimensionless by dividing the misfit at each station by the seismic depth. The of 0.1 km, and had an RMS difference of  $\pm 4.6$  km. This RMS difference could be error and correlated with their gridded equivalents at 0.89, had a mean difference to assess the quality of the grid. The 39 control depths were assumed to be without locations of seismically determined control depths to perform a statistical comparison anomaly data and digital elevation models. This grid is interpolated to the individual A model of the depth to the Moho was derived from available free-air gravity

#### E.1 Introduction

discarded. agreement indicates that one or both are incorrect, and these comparisons will be derived values and seismically derived values indicates that both are correct. determining agreement. It is assumed that general agreement between Moho model seismic depth estimates, so they will be assumed to be with out error for purposes of and around Greenland. model, comparison is made to depths obtained through seismic surveys for regions in lies and digital elevation models for the terrain. To assess the quality of this Moho Chapter 3, a Moho model was generated using available free-air gravity anoma-No information was provided concerning the quality of these

and differenced to determine basic statistics were interpolated from final Moho model to the locations determined for the seismic to constrain the iteration process used to generate the final Moho model. a long wavelength field. gridding the available seismic estimates and smoothing the resulting grid to produce to check the agreement with the final Moho model. The depth model was generated by estimates. These seismic estimates were used to generate a long wavelength depth model and These interpolated values and the seismic estimates are then compared This field was incorporated into the early Moho model Values

were determined by manually interpolating their final maps to obtain estimates of estimates [Chian & Louden[1992; 1994]. For the other surveys, the location data Fechner & Jokat, 1996; Gregersen et al., 1988; Jackson & Reid, 1994; Reid & Jackson, the Moho [Chian There were six separate surveys that provided seismic estimates for the depth to Only two of these surveys provided complete location data with the seismic 8 Louden, 1994; Chian & Louden, 1992; Dahl-Jensen et al., 1998;

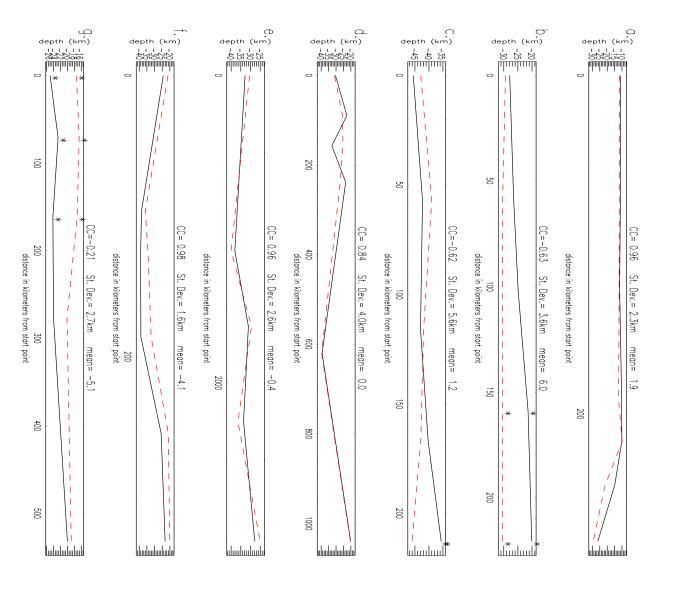
values may occur, which would then affect the statistical comparisons. the depth locations. Due to inaccuracies in this measurement, incorrect interpolated

variations will deviate from the actual Moho depth. not a completely valid assumption. variation. When the Moho model was calculated, a uniform density was assumed for undulations the crustal and mantle material. lation is the chief source of variation in masses and not a lateral or vertical density Another source of errors may derive from the assumption that the Moho undugenerated gravity field, but local density variations will be modeled as Moho instead. Therefore, predictions in regions containing significant density This generated a fixed density contrast. To a first order, this assumption may account

### Comparison of Interpolated Moho Model and Depths Seismic

further in Table E.2. profiles. These profiles are displayed in Figure E.1 and the data points are analyzed stations fall within the actual prediction area, and these fall roughly along 7 different literature sources and interpolated from the Moho shown in Figure Data in Table E.1 show Moho control values (Figure E.3) derived from various  $\mathbb{E}.2$ Only 39

The sixth profile [Jackson and Reid, 1994] and the seventh profile [Reid and Jackson profile [Dahl-Jensen et al., 1998] is over oceanic crust offshore southeastern Greenland. profile [Chian and Louden, 1992] parallels the coast over continental crust. transitional crust and onto continental crust [Chian and Louden, 1994]. The second others are anti-correlative. The fourth profile [Fechner and Jokat, 1996] is over continental crust in the Scoresby From Figure The data from Gregersen et al. E.1, several of the profiles can be seen to agree very The first profile starts over oceanic bedrock passes over [1988] is actually a series of distant points. well, while



removal. and Reid [1994] g. Reid and Jackson [1997]. Jensen et al. [1998] d. Fechner and Jokat [1996] e. Gregersen et al. [1988] f. Jackson each profile. standard deviation (St. seismic estimates given in the red dashed profiles. from the final Moho model (Figure E.2) are shown as solid white profiles with the Figure E.1: Comparison of 7 seismic profiles around Greenland. Values interpolated a. Chian and Louden [1994] b. Dev.), and mean difference (mean) are given at the top of Asterisks (\*) mark points flagged for Chian and Louden [1992] c. The correlation coefficient (CC), Dahl-

and continental crust. 1997] are over northern Baffin Bay in a region that is partially oceanic, transitional,

dependence between crustal type and either under- or over-predicting depths and sixth (oceanic crust) have negative mean values, demonstrating that there crust, and the sixth covers exclusively oceanic crust. Both the fifth (continental crust) first profile covers all types of crust, the fourth and fifth are exclusively continental standard deviations between their interpolated and seismic depth estimates. The first, fourth, fifth, and sixth pairs of profiles show high correlation and low The

further divided by the seismic depth to provide percentage estimates of the magnitude between the control and predicted values are given in kilometers. the amount of disagreement between the interpolated and seismic values is greatest. of the numeric error These data are examined further in Table E.2. In Table E.2, the numeric difference Finally, there are six point flagged with asterisks (\*) indicating points These values are

and three other stations (35, 36, and 37) fell outside of  $\pm 2\sigma$  for the percent RMS outside of the range of  $\pm 2\sigma$  for the numeric differences (stations 11, 12, and 17), difference was 0.1 km, and the percent RMS difference was 18.3%. Three values fell dance with Davis [1986]), the numeric RMS difference was 4.6 km, the mean numeric If no data are removed, then the 39 points correlate at 0.89 (calculated in accor-

the mean numeric difference to -0.5 km, and reduces the percent RMS the CC only to 0.91, reduces the numeric RMS difference only to 4.3 km, increases 17.1%. Removal of the other three stations (denoted with a ‡ in Table E.2) increases numeric difference to 0.9 km, and reduces the percent RMS difference slightly to the CC to 0.93, reduces the numeric RMS difference to 3.7 km, increases the mean Removing only the first group of three (denoted with a † in Table E.2) increases difference

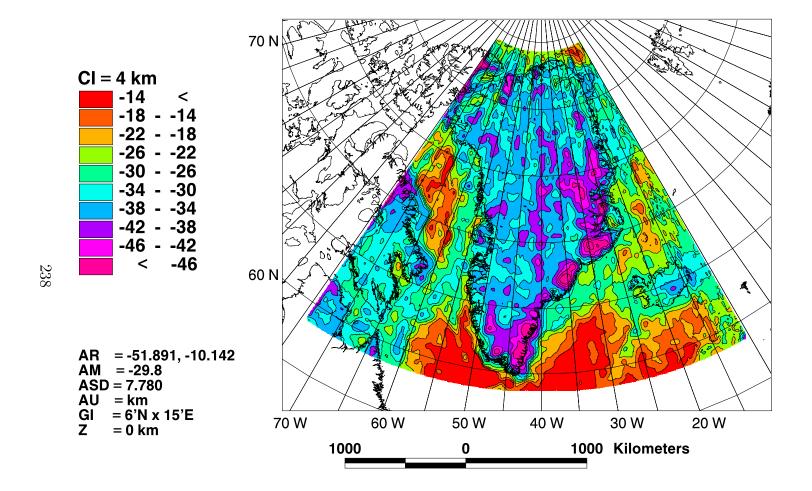


Figure E.2: Moho model for the Greenland field area shown in a Lambert Equal-Area Azimuthal Projection centered on 40 W.

Jo.1	Location	Moho	Depth	sta.	seismic	point
latitude	longitude		mod.	id.	literature	name
(deg. N)	(deg. W)	$(\mathrm{km})$	(km)	no.	source	
59.8408	52.033	11.8	10.9	$\vdash$	Chian & Louden, 1994	R2,A
60.0517	51.2900	11.8	10.9	2	Chian & Louden, 1994	R2,B
60.2767	50.4233	11.9	11.6	ယ	Chian & Louden, 1994	R2,C
60.4442	49.7767	12.3	10.2	4	Chian & Louden, 1994	R2,D
60.5516	49.3514	9.8	10.1	တ	Chian & Louden, 1994	R2,E
60.6633	48.9100	21.4	14.9	6	Chian & Louden, 1994	R2,F
60.7400	48.6133	29.5	26.4	7	Chian & Louden, 1994	R2,G
60.5376	48.2981	29.4	27.9	$\infty$	Chian & Louden, 1992	R1,J
61.0118	49.0482	30.4	26.5	9	Chian & Louden, 1992	R1,K
61.3322	49.5552	30.4	24.9	10	Chian & Louden, 1992	R1,L
61.5943	49.9990	30.4	21.4	11	Chian & Louden, 1992	R1,M
61.8845	50.4750	30.4	20.1	12	Chian & Louden, 1992	R1,N
60.7542	42.4839	42.9	45.6	13	Dahl-Jensen et al., 1998	$50 \mathrm{km}$
61.1960	42.2179	39.0	42.3	14	Dahl-Jensen et al., 1998	$100\mathrm{km}$
61.6007	41.7633	41.9	43.0	15	Dahl-Jensen et al., 1998	$150\mathrm{km}$
62.0555	41.6283	42.9	40.3	16	Dahl-Jensen et al., 1998	$200\mathrm{km}$
62.4576	41.1833	46.0	35.6	17	Dahl-Jensen et al., 1998	$250\mathrm{km}$
70.7689	25.2885	31.0	30.4	18	Fechner & Jokat, 1996	$90544~\mathrm{W}$
70.8297	24.2372	25.7	22.7	19	Fechner & Jokat, 1996	$90544~\mathrm{E}$
70.2230	24.2964	25.3	32.8	20	Fechner & Jokat, 1996	90539 S
70.9568	24.4240	27.2	23.5	21	Fechner & Jokat, 1996	90539 N
70.3933	27.8433	39.7	39.4	22	Fechner & Jokat, 1996	$90320~\mathrm{W}$
70.6289	24.0444	20.3	20.1	23	Fechner & Jokat, 1996	90320 E
69.2254	53.5769	30.5	32.8	24	Gregersen et al., 1988	GDH
77.6812	38.6207	40.0	38.1	25	Gregersen et al., 1988	ILG
81.7101	15.9596	29.5	31.1	26	Gregersen et al., 1988	NOR
76.5507	18.3951	36.5	33.6	27	Gregersen et al., 1988	DAG
70.2778	21.6667	25.0	27.8	28	Gregersen et al., 1988	KTG

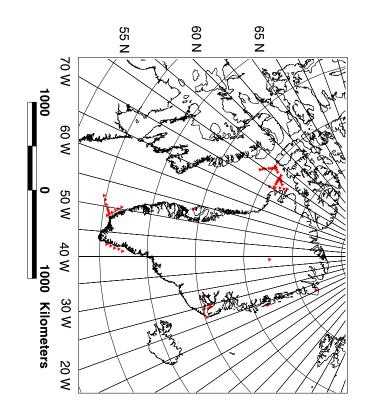
seismic depth values. A dash (-) in the interpolated Moho depth column indicates depth in kilometers interpolated from the Moho model, the seismic literature source, survey literature. Listed are the positional coordinates, seismic depth in kilometers, the values that are analyzed in Table E.2. that have depth values are numerically identified below (stat. id. no.) and map to that the source was outside of the region covered by the Moho grid. The 39 stations and the literature sources' seismic survey line and point names associated with the Table E.1: Seismic control depths for the Greenland Moho model derived from seismic

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4,1	1881	Reid & Jackson,	9	20.0	10.9	6800.67	(9.9949
/ T	1007	0- Tool-gon	၁ ဂ	<b>3</b> 0 0	100	2000 64	79 00 40
4.B	1997	Reid & Jackson.		ı	19.0	74.5146	74.1959
4,I	1997	Reid & Jackson,		ı	19.0	75.1650	74.4148
4,D	1997	Reid & Jackson,		i	19.0	76.0952	74.6819
$_{4,\mathrm{E}}$	1997	Reid & Jackson,		1	19.2	76.5857	74.8974
4,J	1997	Reid & Jackson,		ı	21.3	77.1165	75.0154
$_{4,R}$	1997	Reid & Jackson,		ı	25.4	77.6040	75.1795
$_{4,S}$	1997	Reid & Jackson,		ı	28.0	77.9158	75.2051
$_{2,D}$	1997	Reid & Jackson,	38	24.0	20.1	70.8842	76.1790
$_{2,L}$	1997	Reid & Jackson,	37	24.2	17.2	72.4000	76.1704
$_{2,J}$	1997	Reid & Jackson,	36	22.7	16.6	73.2579	76.1615
$_{2,I}$	1997	Reid & Jackson,	35	25.0	17.4	73.8974	76.1590
$_{2,B}$	1997	Reid & Jackson, 1997		ı	19.3	74.3263	76.1564
$_{2,\mathrm{E}}$	1997	Reid & Jackson,		ı	20.6	75.1111	76.1540
$_{2,S}$	1997	Reid & Jackson,		ı	21.0	75.6211	76.1538
$_{2,R}$	1997	Reid & Jackson,		1	21.0	76.7158	76.1282
$_{3,\mathrm{N}}$	1994	Jackson & Reid,	34	23.1	20.1	70.2991	75.7625
3,G	1994	Jackson & Reid,	33	25.7	21.0	71.0600	76.3052
3,F	1994	Jackson & Reid,	32	38.9	32.2	71.9394	76.7465
3,C	1994	Jackson & Reid,	31	39.1	34.2	72.6122	77.1455
3,A	1994	Jackson & Reid,	30	38.5	36.0	73.2414	77.4225
1,S	1994	Jackson & Reid,	29	24.3	21.2	73.8200	76.7512
1,L	1994	Jackson & Reid,		1	21.3	74.5000	76.5070
1,J	1994	Jackson & Reid,		1	21.3	75.1961	76.2676
1,I	1994	Jackson & Reid,		1	27.4	76.9143	75.6334
1,E	1994	Jackson & Reid,		1	27.6	77.4545	75.3842
1,B	1994	Jackson & Reid,		-	27.6	78.1818	75.0704
		source	no.	(km)	(km)	(deg. W)	$(\deg. N)$
name	e.	literature	id.	mod.	seis.	longitude	latitude
point	,	seismic	sta.	${ m Moho~Depth}$	Moho	Location	Loc
			İ			labie E.I: Continued	Table E.I.

29	28	27	26	25	24	23	22	21	20	19	18	17	16	15	14	13	12	1	10	9	∞	7	6	<b>U</b> T	4	ဃ	2	1		id no.	station
24.3	27.8	33.6	31.1	38.1	32.8	20.1	39.4	23.5	32.8	22.7	30.4	35.6	40.3	43.0	42.3	45.6	20.1	21.4	24.9	26.5	27.9	26.4	14.9	10.1	10.2	11.6	10.9	10.9	$(\mathrm{km})$	predicted	DEPT
21.2	25.0	36.5	29.5	40.0	30.5	20.3	39.7	27.2	25.3	25.7	31.0	46.0	42.9	41.9	39.0	42.9	30.4	30.4	30.4	30.4	29.4	29.5	21.4	9.8	12.3	11.9	11.8	11.8	$(\mathrm{km})$	control	SHI
3.1	2.8	-2.9	1.6	-1.9	2.3	-0.2	-0.3	-3.7	7.5	-3.0	-0.6	-10.4 †	-2.6	1.1	ა. ა.	2.7	-10.3 †	-9.0 +	-5.5	-3.9	-1.5	-3.1	-6.5	0.3	-2.1	-0.3	-0.9	-0.9	$(\mathrm{km})$	num.	DIFFERENCES
14.6	11.2	-7.9	5.4	-4.8	7.5	-1.0	-0.8	-13.6	29.6	-11.7	-1.9	-22.6	-6.1	2.6	8.5	6.3	-33.9	-29.6	-18.1	-12.8	-5.1	-10.5	-30.4	3.1	-17.1	-2.5	-7.6	-7.6	(%)	perc.	ENCES

mic), and percentile difference (the raw difference divided by the seismic depth). depths are given at the bottom of the table. Statistical comparisons excluding no depths, † depths, ‡ depths, and both † and ‡ percentile difference (num. diff.  $\div$  seismic depth) was greater than  $\pm 2\sigma$  (‡) are noted. Depths where the numeric difference was greater than  $\pm 2\sigma$  (†) and depths where the mined Moho depth, the numeric difference of the two depths (predicted minus seis-Columns indicate station identification, predicted Moho depth, seismically deter-Table E.2: Statistical Comparison of the Predicted and Seismic Moho Depths.



in a Lambert Equal-Area Azimuthal Projection centered on 40 W. Figure E.3: Location of Greenland control Moho depths listed in Table E.1 and shown

	Table E.2:
מזיחתות	Continued

13.1	3.4	0.4	0.95	† and ‡
	4.3	-0.5	0.91	† only
	3.7	0.9	0.93	† only
	4.6	0.1	0.89	none
num. RMS	num.	mean		removed
differences	diff		CC	data
ARIS	OMPA	TICAL C	OVERALL STATISTICAL COMPARISONS:	OVERA
	1.1	18.9	20.0	39
	3.9	20.1	24.0	38
	7.0	17.2	24.2	37
	6.1	16.6	22.7	36
	7.6	17.4	25.0	35
	3.0	20.1	23.1	34
	4.7	21.0	25.7	33
	6.7	32.2	38.9	32
	4.9	34.2	39.1	31
	2.5	36.0	38.5	30
	(km)	(km)	$(\mathrm{km})$	
	num.	$\operatorname{control}$	predicted	id no.
DIFFERENCES	DIF	SH	DEPTHS	$\operatorname{station}$
		d	2: Continued	Table E.2:

reduces the percent RMS difference to 13.1%RMS difference to 3.4 km, increases the mean numeric difference only to 0.4 km, and to 15.1%. Removal of all 6 stations increases the CC to 0.95, reduces the

## E.3 Significance of the quality of the predictions

regions between the interpolated and seismic values may arise from errors along the southwestern coastal area and in northern Baffin Bay. The disagreement significantly improved the overall comparisons. These data are primarily located values. The removal of these 6 stations (representing about 15% of the total control) Agreement is taken as an indicator that the Moho model is valid for the other in either or

produce less desirable results. away from the assumption that variations in the Moho depth provide the determined by the seismic profiles. If so, then this would be a significant deviation parison with a nearby profile and that it was anomalously high density material as it to possible serpentinized mantle generated during the rifting of Greenland from assert that a high density feature exists at this point in their profile and attribute ure E.1.b (corresponding to points 11 and 12 in Table E.2). Chian and Louden [1992] intersects the rightmost points in Chian and Louden [1992] survey line shown in Figfeature extends throughout the region along Greenland's southwestern coast, which errors in the control data. Examination of Figure E.2 shows that a very pronounced pear to more a function of the underlying assumptions than the estimation process or The excessive errors located areas along the southwestern coast of Greenland apin the crustal thickness. They speculated that the feature may have some lateral extent Hence, predictions based upon this assumption will primary

from the projected maps in which Dahl-Jensen et al. compass and dividers were used to extract the latitude and longitude information points obtained from Dahl-Jensen et al. [1998] were not given explicitly, (as with underlying assumption for point 17 in Table E.2. Additionally, the locations for the pronounced in its sharpness, it also provides a possible deviation away from the value interpolated from the Moho grid (Figure E.2). Estimation errors in extracting this location data may shift the actual comparison Chian and Louden [1992]) but were instead interpolated from their final map. Figure E.1.c at the rightmost point [Dahl-Jensen et al., 1998]. Similarly a more regional effect may also be observed along the southeastern coast [1998] plotted their results. Although less

and do not map directly into geographic coordinates. locations were re-measured, however, the data were presented in projected coordinates for points 35, [1997] were aslo interpolated from maps. Again, errors in the location determination Similarly, the locations assigned to the measurements from Reid and Jackson 36, and 37 may have contributed to an erroneous comparison.

regional analyses depth estimates will be assumed negligible and the Moho model assumed reliable for standard deviation for Figure E.2. Therefore, the effect of the errors on the Moho that the If these suspect values are disregarded, then the remaining comparisons indicate Moho model is reliable to  $\pm 3.4$  km. This is less than half the  $\pm 8.5$ 

#### E.4 Summary

available throughout the Greenland area. While no error information was available and continental regions. assuming an Airy-Heiskanen model of compensation with fixed densities for oceanic Moho undulation grid was predicted using available digital elevation models and Seismic estimates for the depth to the Moho were also

Moho undulation grid is assumed to indicate reliability. with these data, agreement between them and values interpolated off of a predicted

maps. have had there locations inaccurately determined while extracting that data from the the uniform density model assumed for the entire Greenland area. Other points may ized as a possible serpentinized mantle diapir. This represents a deviation away from the data). Some of these stations occurred within a region that has been characterrelated to the seismic estimates. lated depths was more than  $\pm 2\sigma$  (i.e., these points fell out side of 95% of the rest of There were six stations where the disagreement between the seismic and interpo-Therefore, values interpolated from the final Moho grid may not be correctly

interpolated values implies that the predicted grid is reliable for regional analyses. dard deviation for the Moho model. The overall agreement between the seismic and error of 13%. doubt, the overall agreement is 95% with agreement at  $\pm 3.4$  km or a dimensionless Discounting the six stations whose locations or underlying assumptions are in The  $\pm 3.4$  km RMS error estimate is less than half the  $\pm 8.5$  km stan-

#### APPENDIX F

## GLQ Geophysical Relationships

A density contrast for the prism is also assumed. relationships and positions are fixed by assigning the locations of the nodes contained tational effect for a prism of rock is well established [von Frese, 1980]. The structural within each prism and the points where the gravity effect of those nodes is desired. The use of Gaussian Legendre Quadrature integration (GLQ) to model the gravi-

the prism and its mass. indicates that there are two variables that are relevant to the nodes: apriori knowledge of these positions. However, performing an inversion based upon GLQ is complicated by the lack of Newton's law of gravitation (Equation F.1) the volume of

$$F = -GM_e \frac{m_{prism}}{l^2} \tag{F.1}$$

where: F = the force of gravitational attraction,  $GM_e$  = the gravitational constant times the Earth's mass,  $m_{prism} = \Delta \rho \cdot \text{Volume} = \Delta \rho \cdot \text{height-width-length}$ , and l = the distance between the source and the observation point

assumed density contrast to generate the gravity values at the observation grid in to each prism height to generate a prism that will contain sufficient material at the are fixed, as is the density contrast  $(\Delta \rho)$ . length and width are determined by the selected latitude and longitude spacing and In an inversion, an initial value (1 km) is assumed for the height of the prism. The A scale factor is determined and applied

and, thereby, mass are all linear. The use of a least squares solution necessitates this linear relationship. least squares sense. These relationships between the scale factor, height,

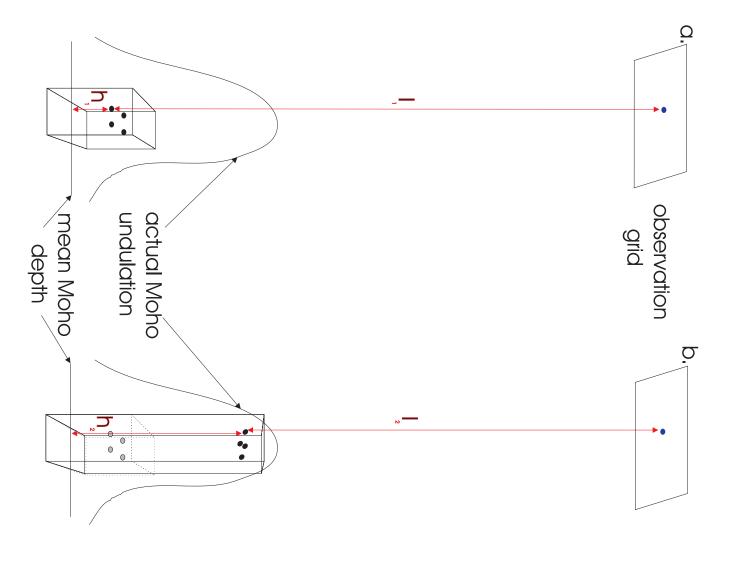
distance between source and observation grids. This is a primary reason for selecting the change in node location due to scaling is sufficiently small when compared to the between the source and observation points (l). be struck. sufficiently high enough elevation to upward continue free-air gravity anomalies. However, changing the height value of the grid also changes the physical distance trade-off is that the signal attenuates with elevation. The linear solution can be applied if Therefore a balance must

and the effect of this secondary relationship diminishes rapidly. solution is convergent). Therefore, the subsequent node location changes are smaller, solution. Because the actual relationship is not linear, iterations are required to generate a In each successive iteration, the scale factor should become smaller (if the

is exemplifying the two relationships between the location and the gravity effect in Equation F.1 and not demonstrating the application of GLQ theory immediately nadir to the the observation point is simplified example is given in Figure F.1. The simple case where the given, as the primary interest node

iterations to generate a complete solution. This secondary change caused by the shift in node location is what necessitates volume As can be seen in Figure F.1, changing nodal height by scaling affects both the of the prism (and, hence, mass) and the distance to the observation

Moho undulation. contrast at the Moho, the Moho undulation itself, and the density contrast at the annihilating compensated terrain gravity effect (ACTGE) generated by the density There are three factors that can be estimated using GLQ: the gravity effect of the Use of GLQ to estimate the ACTGE can be represented



nodal location only for a point immediately above the node in the observation grid. of the prism and its nodes. This simplified case depicts the effects on the change of the prism, its nodes (black dots), and the observation surface. b. The final location and the observation points may be linearly modeled. If  $h_2 - h_1$  is sufficiently small, then  $l_1 \approx l_2$  and the relationship between the prism Figure F.1: Depiction of GLQ Geophysical Relationships. a. The initial location of

the densities are fixed (Equation F.2). used to estimate the unknown ACTGE values (y) through a design matrix (A) where forward solution, where the Moho undulations are known parameters (x) and are

$$y = Ax (F.2)$$

tion (Equation F.3) using the same fixed densities (A). desired parameters  $(\hat{x})$  and the observations (y) are used to generate an inverse solucussed above and represents an inversion where the unknown Moho depths are the Fixing the density contrasts and estimating the Moho from the ACTGE was dis-

$$\hat{x} = (A^t A)^{-1} A^t y \tag{F.3}$$

depths are fixed and density values for each prism are the desired values (z) as shown as the observations (y), but the design matrix is different (B) because the Moho in Equation F.4. The final method for applying GLQ is also an inverse solution using gravity values

$$\hat{z} = (B^t B)^{-1} B^t y \tag{F.4}$$

source of gravity field variations and density variations are only of secondary concern. of the ACTGE but assumes that the effects of the Moho undulation are the primary If this assumption is invalid, then the resulting predictions may be in error. modify initial densities using Equation F.4. This permits a more complete modeling tion F.3, this method may be now applied to model any residual ACTGE values to Once the best estimate for the Moho undulation has been determined from Equa-

contrasts were invalid, the initial models may be updated and new results generated However, if later analysis discovers that the initial assumptions about the density

the constraining data be added in accordance with Equation F.5. incorporated into a solution as constraints  $(\check{x}_2)$ . A revised model would require that process. Likewise if reliable depth estimates are available these data may also be density in a region become available, these data may be incorporated into the GLQ for the Moho depths. If from the literature or data collection, revised estimates for

where: 
$$\hat{x_1} = \text{the remaining unknown depths}$$

$$N_{11}^{-1} = \text{the inversion matrix } (A_1^t A_1)^{-1})$$

$$A = \text{the design matrix and its two components } [A_1 A_2]$$

$$y = \text{the observation vector}$$

$$\hat{x_2} = \text{the constraints}$$
(F.5)

must be regenerated. Because the new data used as constraints would be replacing old cells, the data

#### APPENDIX G

# Kinematic Model Summary of Events for Greenland

regions in terms of major geologic events for the past 130 Ma that were adapted from the Labrador Sea. Table G.1 summarizes the chronological development of these four and the borderlands surrounding it into 6 major areas when they developed their Table 3 of Srivastava and Tapscott [1986]. kinematic model for its opening. Four of these six are relevant to the formation of Srivastava and Tapscott [1986] divided the entire North Atlantic Ocean region

rifting resulted in the separation of Greenland from North America as Greenland tiated within the Labrador Sea as rifting halted in the Rockall-Hutton Trough. This 1993; Dahl-Jensen et al., 1998; Vogt, 1986]. At about 84 Ma, seafloor spreading inithat was a part of the South Greenland Archean province [Toft and Arkani-Hamed, of Greenland. The Rockall-Hutton Bank represented a piece of Precambrian terrain east of the Rockall-Hutton Bank, which was then located off of the southeastern coast of Laurasia and Gondwana. Spreading initially took place in the Rockall Trough just lantic started approximately 95 Ma, progressing from the south along the separation began to move with Europe. According to Srivastava and Tapscott [1986], the initial rifting of the North At-

and initiated compression in the northwest with Ellesmere Island in what would Seafloor spreading in southwestern Greenland caused a counter-clockwise rotation

the Eurasian continental shelf to the north. This change in direction coincided with the separation of the Lomonosov Ridge from now mostly N-S as Greenland separated from the Rockall-Hutton Bank and Europe. (ca. 56 to 50 Ma), a 45° shift in spreading direction occurred, such that motion was 25 (ca. 84 to 60 Ma) was largely in an ENE direction. From isochrons 24 through 21 become the Eurekan Orogeny. Initial motion of Greenland from isochrons 33 through

halted. From that time to the present, Greenland has been relatively quiescent. At that time Greenland began to move with North America and the Eurekan Orogeny Eventually, the Labrador spreading ridge failed at about isochron 13 (ca. 20 Ma).

Anomaly	Age (Ma)	Arctic Basin	Greenland- Norwegian Seas
M11	Hauterivian (130)	Opening of the Canada Basin	
M0	Aptian (118)	Near extinction of spreading in Canada Basin.	
M0-34	Cenomanian (95)	Initiation of compression between Lomonosov Ridge and on Siberian Platform possibly along Verkhoyansk fold belt.	
34	Campanian (84)	Some spreading in the Makarov Basin. Compression along Verkhoyansk fold belt.	Rifting in the Greenland-Norwegian Sea.
31	Maastrichtian (68)	Some spreading in the Makarov Basin.	As above.
25	Thanetian (59)	Shearing in the northern Siberian Platform due to separation of Lomonosov Ridge from the Barents Shelf.	Volcanism in eastern Greenland near Scoresby Sund and on the Voring Plateau, Faeroe Islands and initiation of Greenland-Scotland Ridge.
24	Ypresian (56) (E. Eocene)	Dilation in the northern part of the Siberian Platform and compression in the southern part, possibly along Verkhoyansk fold belt.	Active sea-floor spreading in Norwegian and Greenland Seas.
21	Lutetian (50) (M. Eocene)	As above.	Formation of a large part of Iceland-Faeroe Ridge. Possible jump of ridge axis to the west.
13	Rupelian (36) (E. Oligocene)	Shearing motion on Siberian Platform along Leona River.	Sea-floor spreading in Norwegian and Greenland Seas continuing. Simultaneous spreading east and west of Jan Mayan Ridge.
7	Chattian (26) (L. Oligocene)	As above.	Jan Mayan Ridge separates from Greenland and spreading starts Kol- beinsey Ridge. Spreading in the Nor- wegian Sea and on the Iceland-Faeroe Ridge terminates and starts in Iceland.
0	Present	Strike-slip motion along the intercontinental boundary between Eurasia and North America on the Siberian Platform.	Spreading taking place across Reykjanes, Kolbeinsey, Mohn and Knipovich Ridges and in Iceland.

adapted from Table 3 of Srivastava and Tapscott [1986]

Continued

Table G.1: Chronological summary of tectonic events for the Greenland area. The conventional name of the associated magnetic isochron is given under the "anomaly" events of their evolution. heading. Geologic age and associated time in millions of years is given under the "age (Ma)" heading. Subsequent columns for the listed regions summarize the geologic

Anomaly M11	Anomaly Age (Ma) I	Q Labrador Sea 	North Atlantic
MII	Hauterivian (130)		
M0	Aptian (118)	Volcanism on the southern Labrador Shelf and on land associated with the initial stages of sea-floor spreading in the Labrador Sea.	Initiation of rifting between the Porcupine Bank and Orphan Knoll and in the Rockall-Hutton Trough.
M0-34	Cenomanian (95)	Rifting between Greenland and Labrador, volcanism on the outer shelf.	Initiation of active sea-floor spreading in the North Atlantic and Rockall-Hutton Trough (all regions south of the Charlie Gibbs FZ).
34	Campanian (84)	Active sea-floor spreading in the south Labrador sea and rifting in the north Labrador Sea and Baffin Bay. Counterclockwise rotation of Greenland relative to Ellesmere Island and start of compression in the northern Sverdrup Basin.	Active sea-floor spreading in all regions south of Greenland, jump of ridge axis to the west and spreading stopped in Rockall-Hutton Trough.
31	Maastrichtian (68)	Active sea-floor spreading in the north-Labrador Sea and rifting in the northern Baffin Bay. Compression in the Sverdrup Basin and start of the Eurekan Orogeny.	Active sea-floor spreading continuing.
25	Thanetian (59)	Volcanism in the Davis Strait and surrounding regions, change in direction of motion between Greenland and North America; Greenland begins to move at an angle to the Nares Strait.	Volcanism forming the Thulean Rise. Formation of the Azores-Biscay Rise.
24	Ypresian (56) (E. Eocene)	Change in direction of motion between Greenland and North America. Oblique spreading in Baffin Bay and Davis Strait resulting in mainly shear motion.	Compression and strike-slip movement along the Iberia-Africa and Eurasian boundary.
21	Lutetian (50) (M. Eocene)	Change in direction of motion between Greenland and North America. Motion between Greenland and Ellesmere Island is mainly compressional.	As above and start of formation of King's Trough.
13	Rupelian (36) (E. Oligocene)	Extinction of sea-floor spreading in Baffin Bay and Labrador Sea. Greenland started to move with North America.	Motion along King's Trough continuing.
7	Chattian (26) (L. Oligocene)	1	Motion along King's Trough stopped. Plate boundary between Eurasia and Africa shifts to the south along the Azores-Gibraltar FZ.
0	Present	1	Spreading taking place across the MAR.

#### APPENDIX H

## Acronym Definitions

ACTGE Annihilating Compensated Terrain Gravity Effect

AGC Atlantic Geoscience Centre

ASDStandard Deviation

BRTBottom Relief Topography

CTGE Compensated Terrain Gravity Effect

DEM Digital Elevation Model

DLFIDifferenced Local Favorability Indices

FAGA Free-Air Gravity Anomalies

FVD(CGE) First Vertical Derivative of Crustal Gravity Effect

FVD(IC-TDFAGA)First Vertical Derivative of IC-TDFAGA

GAP91 GAP92Greenland Aerogeophysical Project 1991

Greenland Aerogeophysical Project 1992

GGMGravity Geologic Method

GLQGaussian Legendre Quadrature integration

GQLFI Gravity Quotient Local Favorability Indices

GSC Geological Survey of Canada

IC-RTPMA IntraCrustal Reduced-To-Pole Magnetic Anomalies

IC-TDFAGA IntraCrustal Terrain Decorrelated Free-Air Gravity Anomalies

IGRF International Geomagnetic Reference Model

LH Local Favorability Indices

MA Magnetic Anomalies

MC-RTPMA Mantle/Core Reduced-To-Pole Magnetic Anomalies

MQLFI Magnetic Quotient Local Favorability Indices

MSLMean Sea Level

NRLNaval Research Laboratory

OBSOBServed

REGQLFI REGional Quotient LFI

RES RESidual

RTPRTPMA Reduced-To-Pole Magnetic Anomalies Reduced-To-Pole

SLFISummed Local Favorability Indices

TCFAGATerrain Correlated Free-Air Gravity Anomalies

TDFAGA Terrain Decorrelated Free-Air Gravity Anomalies

TGE Terrain Gravity Effect